

Manuscript Number: LITHOS6792R2

Title: The magmatic evolution and genesis of the Quaternary basanite-trachyphonolite suite of Itasy (Madagascar) as inferred by geochemistry, Sr-Nd-Pb isotopes and trace element distribution in coexisting phases

Article Type: Regular Article

Keywords: basanites, trachyphonolites, mineral trace elements, titanite removal, crustal contamination, Itasy, Madagascar

Corresponding Author: Dr. Leone Melluso,

Corresponding Author's Institution: Università di Napoli Federico II

First Author: Leone Melluso

Order of Authors: Leone Melluso; Tucker D Robert; Ciro Cucciniello; Anton P le Roex; Vincenzo Morra; Rocky L Rakotoson, Dr.

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## Revision Notes

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Dear Editors of Lithos,

in this revised version, we specifically rewrote the last part of the Introduction and Geological setting, from line 97 to line 106, in order to closely meet the requirements of the Reviewer and to not make this introduction too wordy.

We also made minor corrections to the English text throughout, and added an article of ours, now in press with doi. No changes were made to tables, figures and supplementary material.

## Abstract

The Itasy is a Pleistocene-Holocene volcanic field in central Madagascar, located to the west of the Ankaratra volcanic complex. It comprises scoria cones and lava domes (>120), with associated pyroclastic fall and mafic lava flows, covering an area of ab. 400 km<sup>2</sup>. The last volcanic episodes probably dated ca. 6000-7100 y BP; warm springs and geysers are active. The juvenile samples comprise a peculiar, almost bimodal, rock suite ranging from potassic leucite-kaersutite-bearing basanites, tephrites and phonotephrites, to benmoreites and titanite-haüyne-bearing trachyphonolites (MgO from 9-10 wt.% to 0.1 wt.%). These rocks show continuous and overlapping variations in the bulk-rock and phase composition (olivine, clinopyroxene, amphibole, feldspar, leucite, haüyne, nepheline, oxides, apatite, titanite, glass and other accessories). The basanites have homogeneous isotopic composition ( $^{87}\text{Sr}/^{86}\text{Sr}=0.70366-0.70378$ ,  $^{143}\text{Nd}/^{144}\text{Nd}=0.51274-0.51277$ ,  $^{206}\text{Pb}/^{204}\text{Pb}=18.7-18.9$ ,  $^{207}\text{Pb}/^{204}\text{Pb}=15.53-15.56$ ;  $^{208}\text{Pb}/^{204}\text{Pb}=38.89-39.01$ ), and a marked enrichment in the most incompatible elements (LILE and HFSE reach 100-215 times primitive mantle). These features are consistent with low degrees of partial melting of a volatile-, LILE- and HFSE-rich, amphibole-bearing peridotitic mantle under conditions of uplift during an E-W-directed extensional regime, as found in central Madagascar. The marked changes in the geochemical composition, and small variations of the Sr-Nd-Pb isotopes shown by the trachyphonolites ( $^{87}\text{Sr}/^{86}\text{Sr}=0.70425-0.70446$ ,  $^{143}\text{Nd}/^{144}\text{Nd}=0.51266-0.51269$ ,  $^{206}\text{Pb}/^{204}\text{Pb}=18.18-18.39$ ,  $^{207}\text{Pb}/^{204}\text{Pb}=15.49-15.51$ ;  $^{208}\text{Pb}/^{204}\text{Pb}=38.38-39.57$ ) with respect to basanites and tephrites point to a limited amount of crustal contamination of the relatively low- $^{206}\text{Pb}/^{204}\text{Pb}$ , low- $^{143}\text{Nd}/^{144}\text{Nd}$ , high- $^{87}\text{Sr}/^{86}\text{Sr}$  Precambrian basement rocks (of Middle Archean to Late Proterozoic age), and highlight the geochemical effects of titanite and anorthoclase removal on the trace element fractionation trends, as also shown in the trace element composition of the phenocrysts in the trachyphonolites.

The Itasy volcanic field has basanites, tephrites, phonotephrites, benmoreites and trachyphonolites

These rocks are the result of fractional crystallization and crustal contamination

Titanite and anorthoclase removal is essential petrogenetic process in the trachyphonolites

The mantle source of the basanites is similar to that of the other Cenozoic rocks of central Madagascar

1 **The magmatic evolution and genesis of the Quaternary basanite-trachyphonolite suite of Itasy**  
2 **(Madagascar) as inferred by geochemistry, Sr-Nd-Pb isotopes and trace element distribution**  
3 **in coexisting phases**

4  
5 L. Melluso\*<sup>1</sup>, R.D. Tucker\*\*, C. Cucciniello\*, A.P. le Roex\*\*\*, V. Morra\*, A. Zanetti\*\*\*\*, R.L.  
6 Rakotoson\*\*\*\*\*

7  
8 \*DISTAR, Università di Napoli Federico II, Napoli, Italy

9 \*\*BRGM, Orleans, France

10 \*\*\*University of Cape Town, Rondebosch, Republic of South Africa

11 \*\*\*\*IGG-CNR, Pavia, Italy

12 \*\*\*\*\*Analamaity, Antananarivo, Madagascar

13  
14 <sup>1</sup>corresponding author: melluso@unina.it

15  
16 **Abstract**

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40

41 Keywords: basanites, trachyphonolites, mineral trace elements, titanite removal, crustal  
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43

## 44 **1. Introduction and Geological setting**

45

46 Madagascar is the site of abundant Cenozoic-to-Recent volcanism (Woolley, 2001; Tucker et al.,  
47 2008; Cucciniello et al., 2011, 2016, 2017), which crops out from the northernmost regions of the  
48 island (Cap d'Ambre/Bobaomby, Massif d'Ambre) to the southwest (Ankililoaka, Tulear region;  
49 Fig. 1a). This volcanism is the response to extensional processes and uplift active since the Cenozoic  
50 of the Madagascan lithosphere, with changing orientation of the stress pattern in different sectors of  
51 the island (Nicollet, 1984; Piqué et al., 1999; Cucciniello et al., 2016, 2018). The changing ages and  
52 lithological characteristics of the exposed crustal domains are reflected in the different isotopic  
53 composition of the mantle-derived Cenozoic rocks, suggesting a strong lithospheric control in the  
54 genesis of the latter (Tucker et al., 2014; Melluso et al., 2016; Cucciniello et al., 2018).

55 The central field of alkaline igneous rocks consist of four different volcanic-dominated massifs,  
56 from east to west: Takarindiona, Alaotra, Ankaratra and Itasy (Fig. 1a). The eastern massifs of  
57 Takarindiona and Alaotra are relatively small, structurally controlled Miocene-Pliocene fissural  
58 eruptions and spatter cones (Melluso et al., 2011a), emplaced along or near the western margin of  
59 the oldest rocks in Madagascar (~3.2 Ga gneisses of the Antongil/Masora domain). The Ankaratra  
60 and Itasy are significantly larger in size and more diverse in morphology, as they include lava  
61 domes, massive flow eruptions, spatter cones, and various phreatomagmatic rings and cones (cf.  
62 Brenon and Bussiere, 1959; Woolley, 2001; Cuciniello et al., 2017). The Ankaratra covers an area of  
63 ~3600 km<sup>2</sup> and reaches 2395 mt. elevation. Like the eastern fields, the age of activity in the  
64 Ankaratra spans from Miocene to Recent (Tucker et al., 2008; Rufer et al., 2014; Cucciniello et al.,  
65 2017).

66 The Itasy volcanic field is a much smaller (roughly 400 km<sup>2</sup>) and younger version of the Ankaratra;  
67 it crops out at roughly 1200 mt. a.s.l., in a north-south-trending graben located in the Precambrian  
68 basement (e.g., Andrianaivo and Ramasianoro, 2010, 2011), whose western limits are best indicated  
69 by a long NNE-SSW normal fault (Fig. 1b). The volcanic field is the coalescence of more than 120  
70 scoria cones (e.g., Kassigie, also known for a presumed eruption in 2001, likely a landslide in its SW

71 sector), with lava flows (e.g., along the western master fault), lava domes (e.g., Ngilofotsy, west of  
72 Analavory; Fig. 1b; Supplementary Fig. 1), and pyroclastic rocks in the facies of pyroclastic fall  
73 deposits, small-scale pyroclastic flows and maar craters (cf. Supplementary Fig. 1; Razafiniparany et  
74 al., 1974). According to Bussiere (1957, and references therein), the sequence of volcanic events is  
75 established by cross-cutting relationships. It includes: (a) oldest eruptions of trachyte and  
76 trachyphonolite, (b) deposition of lacustrine sediments bearing fossils of Pleistocene or younger age  
77 (i.e. plant fragments, and bones of hippos, lemurs, and *Aepyornis*), (c) lava flows and pyroclastic  
78 eruptions of basanite and tephrite, and (d) recent strombolian and phreatomagmatic eruptions of  
79 basanitoids. Given the small volume, the scattered nature of magmatism and the structural  
80 complexity of this part of Madagascar, the eruptive history of the Itasy area is only generally known.  
81 What is clear, however, is that the age of the volcanism must be young, certainly less than 1 Ma.  
82 This is consistent with the geologic youthfulness of the volcanic landforms, their relationship to  
83 Pleistocene or younger fossils, the historical record of seismic and geothermal activity (Brown et al.,  
84 2015), and an  $^{40}\text{Ar}/^{39}\text{Ar}$  plateau age of  $97\pm 6$  ka for a trachyphonolite sampled near Analavory (R.D.  
85 Tucker, unpub. data). During and after formation of the lacustrine basins, the region is believed to  
86 have experienced regional E-W extension and differential uplift that fractured the natural dams,  
87 drained the tectonic lakes, and produced its rugged topography. The bounding faults of the Itasy  
88 graben (Fig. 1b) are but one manifestation of this near-surface extensional stress field. These  
89 features, including the volcanism of Itasy and Ankaratra, are a response to still unresolved forces,  
90 recently imaged in a regional seismological survey (Pratt et al., 2017; Fig. 1c) that acted throughout  
91 the Malagasy lithosphere in Miocene and younger time.

92 The Itasy and the Ankaratra are built upon a Precambrian basement complex of Neoproterozoic granitic  
93 gneiss (2.7-2.5 Ga, Betsiboka Suite) and Neoproterozoic quartzite and schist (<1.0 Ga,  
94 Ambatolampy Group), and both are bounded by faults near the apex of the Taná virgation, where the  
95 gneissic fabrics turn sharply from N-S to E-W strike directions (Roig et al., 2012; Tucker et al.,  
96 2014).

97 Petrogenetic studies on the Ankaratra volcanic rocks were recently published (Melluso et al., 2016;  
98 Cucciniello et al., 2017). This paper is focused on the petrogenesis of the Itasy, probably the least  
99 known Late Cenozoic *volcanic* field in Madagascar. Lacroix (1923) and Bussiere (1957) first  
100 identified the featuring petrographic characteristics of the main lithologies, including the presence of  
101 h  y  ne in the felsic samples, allowing them to make comparison to rocks of the Massif Central in  
102 France. We report the range of chemical and isotopic composition of the Itasy bulk rocks and  
103 mineral phases, and highlight magmatic processes of within-plate magma series which acted during  
104 the ascent in the crust. We also focus on the petrogenetic relationships and differences with other



105 Cenozoic volcanic districts of Madagascar, particularly those of the central Madagascar, which were  
106 emplaced through the same Precambrian domains.

107

## 108 **2. Classification and petrography**

109

110 Sixty-nine samples were collected throughout the volcanic field (Fig. 1b). They are representative of  
111 scoria and lava flows cropping out in the graben, and of the several volcanic domes of evolved  
112 composition (trachytic *s.l.*), which occur in the same areas as scoria cones and lavas. The analytical  
113 techniques are fully described in the appendix. According to petrography, T.A.S.,  $R_1R_2$  classification  
114 diagrams (Fig. 2), and CIPW norms (Supplementary Table S1), the fresh samples are basanites,  
115 tephrites, rare phonotephrites and benmoreites (*trachyandesites*), and trachyphonolites, with a  
116 bimodal character (Fig. 2, Table 1 and Supplementary Table S1). The *mafic* rocks have potassic  
117 affinity, given that their  $\text{Na}_2\text{O}/\text{K}_2\text{O}$  ratios are close to the unity; the CIPW norms have no normative  
118 leucite. The trachyphonolites do not reach fully peralkaline conditions [P.I., peralkaline index, molar  
119  $(\text{Na}+\text{K})/\text{Al} < 0.98$ ], with the exception of the groundmass of evolved samples (see below). The  
120 samples with L.O.I. values  $> 2$  wt.% (mostly scoria) have signs of chemical alteration, and thus were  
121 excluded from the discussion and from the classification diagrams. A synopsis of the observed  
122 mineral assemblages for various lithotypes is reported in the Supplementary Table S2.

123 **Basanites, tephrites and phonotephrites** are variably porphyritic (in both lava and scoria facies),  
124 with zoned, Ti-rich purple-green clinopyroxene, olivine (with small Cr-spinel inclusions), amphibole  
125 and oxide phenocrysts and megacrysts in a fine grained, feldspar-rich mesostasis, to almost aphyric  
126 facies, with purple-green clinopyroxene, oxide and olivine microlites in fine-grained or glassy  
127 groundmass, often without amphibole. The latter phase is found in both scorias and lavas  
128 (Supplementary Fig. S1). Amphibole is almost always poikilitic over clinopyroxene (Supplementary  
129 Fig. S1), and may show reaction relationships with the latter; it is often replaced by secondary  
130 rhönite. There is complete lack of feldspar phenocrysts. Apatite microphenocrysts and microlites are  
131 ubiquitous, particularly in the phonotephrite RT-06I-398, and leucite is often present in the  
132 groundmass, together with h aüyne, phonolitic glass and alkali feldspar (Supplementary Fig. S1;  
133 Supplementary Table S2).

134 **Benmoreites**; sample RT06I-387 (the “*ordanchite*” of Lacroix and Bussiere) has rare phenocrysts of  
135 plagioclase, clinopyroxene, amphibole and opaque oxides, set in a trachytic/aphanitic mesostasis  
136 (Supplementary Fig. S1).

137 **Trachyphonolites** range from porphyritic to nearly aphyric, and are characterized by phenocrysts of  
138 anorthoclase and/or sanidine, plus rarer titanite, green clinopyroxene, amphibole and opaque oxides,

139 set in a trachytic mesostasis that also has apatite, nepheline, hauyne, sodalite, clinopyroxene, other  
140 accessories and interstitial glass (Supplementary Fig. S1; Supplementary Table S2).

141

### 142 **3. Mineral compositions**

143

144 **Olivine** in basanites and tephrites is usually zoned, with composition ranging from Fo<sub>81</sub> to Fo<sub>55</sub> [Fo  
145 is 100 Mg/(Mg+Fe) in atoms], with MnO up to 0.52 wt.% and CaO up to 0.73 wt.% in the most Fe-  
146 rich compositions (Supplementary Table S3). Some crystals appear too Mg-rich to be in equilibrium  
147 with the host rock composition [e.g., the rounded Fo<sub>81</sub> (xeno)crystal in the RT-06I-398  
148 phonotephrite; Supplementary Table S3].

149 **Oxides. Cr-spinel** is rare and chemically zoned in the Mg-rich olivine cores of basanite RT-06I-  
150 354B [Cr<sub>2</sub>O<sub>3</sub> =8.7-28.7 wt.%; Cr#=47-57; Cr# is 100Cr/(Cr+Al) in atoms]. **Ilmenite** is also very  
151 rare; it was found in the basanite RT-06I-354b as an exsolved phase (Supplementary Fig. S1) and is  
152 rich in geikielite component (MgO=7.8-13.3 wt.%), consistent with its crystallization from a mafic  
153 melt. **Titaniferous magnetite** is the main oxide; it shows increase in MnO and magnetite  
154 component (*s.s.*) and decrease in TiO<sub>2</sub>, Al<sub>2</sub>O<sub>3</sub> and MgO from mafic to evolved rocks (Supplementary  
155 Table S4 and Supplementary Fig. S2).

156 **Clinopyroxene** ranges from diopside [Mg#=82, TiO<sub>2</sub>=1.3 wt.%; Mg# is 100Mg/(Mg+Fe)] to  
157 titanaugite (Mg#=68; TiO<sub>2</sub>=5.75 wt.%, ca. 19 mol.% CaTiAl<sub>2</sub>O<sub>6</sub>) in basanites and tephrites, and  
158 from titanaugite (Mg#=67; TiO<sub>2</sub>=5.1 wt.%) to aegirinaugite (Mg#=22; Na<sub>2</sub>O=5 wt.%, ca. 38 mol.%  
159 aegirine) in the trachyphonolites (Fig. 3). The composition of clinopyroxene is therefore indicative  
160 of the degree of evolution of the host lithotypes, as shown in the Ti-Al and Mg-Ti diagrams  
161 (Supplementary Fig. S3). The evolved rocks have the lowest Mg#, Ti and Al and the highest Mn  
162 (and Na) concentrations; the calculated Al<sup>VI</sup> is absent or strongly subordinated with respect to Al<sup>IV</sup>  
163 (Supplementary Fig. S3; Supplementary Table S5). The compositional ranges are well within those  
164 already observed in the Ankaratra and in the other Cenozoic volcanic areas of Madagascar (e.g.,  
165 Melluso et al., 2011a; Cucciniello et al., 2017 and references therein).

166 **Amphibole** is potassian kaersutite in basanites and tephrites (TiO<sub>2</sub>=4.1-5.1 wt.%; Mg#=58-72), and  
167 potassian kaersutite/Fe-kaersutite to pargasite (TiO<sub>2</sub>=1.2-4.6 wt.%; Mg#=23-56) in benmoreite and  
168 trachyphonolites (Fig. 3). Amphibole has the highest Mg# in the most Mg-rich lava sample (RT-06I-  
169 354B), and has *systematically* lower Mg# than the coexisting clinopyroxene, confirming its later  
170 appearance in the crystallization order (Supplementary Table S6; Supplementary Fig. S1), and thus  
171 bearing no chemical evidence for its crystallization at mantle depths. As with clinopyroxene, there is  
172 no correlation between proxies of the crystallization pressure (e.g., the calculated Al<sup>VI</sup>) and the  
173 composition of amphibole in its host lithology (Supplementary Fig. S4). **Biotite** (Mg#=70-76;

174 TiO<sub>2</sub>=6.5-6.8 wt.%) is very rare. **Rhönite** [Ca<sub>2</sub>(Mg,Fe<sup>2+</sup>,Fe<sup>3+</sup>,Ti)<sub>6</sub>(Si,Al)<sub>6</sub>O<sub>20</sub>] is often found as a  
175 (subsolidus) rim of amphibole.

176 **Feldspar.** Tiny microlites of labradorite (An<sub>56-59</sub>) to oligoclase (An<sub>27</sub>) are found in the groundmass  
177 of basanites and tephrites, with additional interstitial anorthoclase (Fig. 4a), whereas a continuous  
178 trend from andesine (An<sub>42</sub>) to anorthoclase and sanidine (up to Or<sub>57</sub>) is found in the trachyphonolites  
179 (Fig. 4a). Sr is higher than Ba in plagioclase; they reach 1.9 and 4 wt.%, as oxides, respectively, in  
180 alkali feldspar grains of trachyphonolites (Supplementary Table S7). The overall chemistry is more  
181 restricted than that observed for the feldspars at the Ankaratra volcanic complex (Fig. 4a).

182 The interstitial **glass** is trachyphonolitic to phonolitic in composition (Supplementary Table S7).

183 **Feldspathoids.** **Leucite** has been found in the groundmass of the basanites; **häüyne** is found in  
184 phonotephrites, benmoreites and, particularly, in the groundmass of the trachyphonolites. Häüyne  
185 has a significant compositional range, with an almost continuous increase of Na and Cl and decrease  
186 of Ca, K and S towards the late crystallized **sodalite** microlites, also found as rims of zoned crystals  
187 (Fig. 4b, 4c; Supplementary Fig. S1; Supplementary Table S8). The composition of the volatile-  
188 bearing feldspathoids broadly matches the chemical variation of the Mt. Vulture analogues (which  
189 are likely the most complete series of sodalite-group *magmatic* feldspathoids), with the exclusion of  
190 the most potassic varieties (cf. Fig. 4b, and Melluso et al., 2011b and references therein). A Si-rich  
191 **nepheline** (Q<sub>8-24</sub>K<sub>S5-13</sub>Ne<sub>68-78</sub> in wt.%), with low K<sub>2</sub>O (up to 4 wt.%) and variable CaO (up to 2.8  
192 wt.%) completes the groundmass assemblage of holocrystalline samples, including basanites and  
193 tephrites (Supplementary Fig. S5; Supplementary Table S8). The Itasy nepheline differs most  
194 markedly from that of the Madagascan olivine melilitites, the latter having low silica and very high  
195 K (10-12.6 wt.% K<sub>2</sub>O; Melluso et al., 2011a; Supplementary Fig. 5), though it still occurs as a  
196 groundmass phase, and matches the nephelines of Ankaratra and Ankililoaka (Cucciniello et al.,  
197 2017, 2018).

198 **Titanite** occurs as rare phenocrysts in benmoreite and trachyphonolites (Supplementary Fig. 1) and  
199 modally decreases in the most evolved rocks; it usually co-crystallizes with clinopyroxene. The FeO<sub>t</sub>  
200 varies from 1.6 and 7.4 wt.%, the  $\sum$ REE<sub>2</sub>O<sub>3</sub> range from 1.4 to 5.7 wt.%, and ZrO<sub>2</sub> reaches  
201 concentrations as high as 7.8 wt.%, but it is usually far lower (0-4 wt.%). **Apatite** is ubiquitous as a  
202 groundmass phase, and can be found also as microphenocrysts in the phonotephrites. The  
203 concentration of F varies from 1.4 to 4.9 wt.% Cl can be high (up to 1.8 wt.%; average 1.19 wt.%),  
204 and SO<sub>3</sub> is locally significant (up to 1 wt.% in crystals of a tephrite and 2.3 wt.% in crystals of  
205 phonotephrite RT-06I-398; Supplementary Table S9). Rare crystals of **britholite**  
206 [(Ca,LREE,Th)<sub>5</sub>(SiO<sub>4</sub>,PO<sub>4</sub>)<sub>3</sub>(OH,F)] (SiO<sub>2</sub>=21.7 wt.%;  $\sum$ REE<sub>2</sub>O<sub>3</sub>=52 wt.%, F=1.9 wt.%, no Cl and  
207 SO<sub>3</sub>) are found in the groundmass of trachyphonolites. **Zircon** has been found in the interstices of a  
208 trachyphonolite. **Fluorite** is also noted, as well as **hiortdahlite** [ideally

209  $(\text{Ca,Na})_3(\text{Zr,Ti})\text{Si}_2\text{O}_7(\text{F,O,OH})_2]$  in the groundmass of the trachyphonolite RT-061-379  
210 (Supplementary Table 9). Hiortdahlite has been noted in the Ampasindava phonolites (northern  
211 Madagascar; Cucciniello et al., 2016) and, together with the Na-rich clinopyroxene, indicates the  
212 transition of the groundmass liquids to peralkaline compositions. Other accessories and secondary  
213 zeolites (mostly analcime around nepheline) are present in the interstices.

214

#### 215 **4. Bulk-rock geochemistry**

216

217 The concentration of MgO, Cr and Ni in the most primitive basanites (9-9.2 wt.%, 277-288 ppm and  
218 154-164 ppm, respectively), the Mg# (56-59) and the relatively low forsterite content of the olivine  
219 phenocrysts indicate that the most primitive Itasy rocks can be considered as evolved liquid  
220 compositions which suffered previous removal of Mg-rich olivine, Cr-spinel and Mg-rich  
221 clinopyroxene from more primitive basanites. At the same time, basanites and tephrites have high  
222 concentration of  $\text{TiO}_2$  (3-4.7 wt.%) and  $\text{P}_2\text{O}_5$  (0.6-1.9 wt.%) (Table 1; Supplementary Table S1). The  
223 variations of CaO,  $\text{TiO}_2$ ,  $\text{Al}_2\text{O}_3$ ,  $\text{Fe}_2\text{O}_{3\text{T}}$  and alkalis of the mafic rocks, and the mineral composition  
224 in basanites and tephrites depict trends at least qualitatively compatible with removal of pyroxene,  
225 amphibole, olivine and oxides, and minor or negligible removal of feldspar and apatite  
226 (Supplementary Fig. S5). The transition from phonotephrite to benmoreite and trachyphonolites is  
227 characterized by no significant variation of  $\text{Al}_2\text{O}_3$ , a change of slope of many minor and trace  
228 element trends, a drop in the concentration of  $\text{P}_2\text{O}_5$ , and a decrease of Y with MgO. The  
229 concentration of Ba and Sr also drops in the most evolved trachyphonolites (Supplementary Fig. S6).  
230 The variations of Rb and Ba of the basanites and other features suggest some heterogeneity in the  
231 composition of the mafic parental magmas. The Zr/Hf ratios increase from 45 in the basanites to 60  
232 in the evolved trachyphonolites, whereas the Nb/Ta ratio is 14-15 in the basanites and 21-28 in the  
233 most evolved trachyphonolites (Table 1; Supplementary Fig. S7). The La/Nb, Ba/Nb, Ce/Pb and  
234 Nb/U ratios of the least evolved basanites (La/Nb=0.83-0.88, Ba/Nb=8-12, Ce/Pb=30-41 and  
235 Nb/U=42-52) match the values expected from mantle-derived magmas free of effects related to  
236 crustal contamination in the source or *en route* to the surface, and the low Zr/Nb of the basanites  
237 (4.2-4.8) indicate a highly enriched source region (Melluso et al., 2016 and references therein). The  
238 REE patterns of the mafic rocks (Supplementary Fig. S8a) are smooth and highly fractionated  
239 (La/Yb<sub>n</sub>=24-27), have no peaks or troughs at Eu, and increase in the total REE concentration (279-  
240 800 ppm) from basanites to the phonotephrites. Starting from the phonotephrites RT-06I-398 and  
241 RT-06I-397, the REE patterns of benmoreites and trachyphonolites become increasingly concave,  
242 with a marked decrease of intermediate-REE to values even lower than those of the basanites  
243 (sample RT-061-357B; Supplementary Fig. S8a, c, d, e), largely constant La<sub>n</sub> and Lu<sub>n</sub>, and small

244 troughs at Eu ( $Eu/Eu^*=0.79$ ). These characteristics are highlighted when REE are compared to the  
245 concentration of the least evolved basanite RT-061-354B (Supplementary Fig. S8e). The mantle  
246 normalized patterns of the Itasy basanites (Fig. 5a, 5b) are almost flat from Rb to Th, have a peak at  
247 Nb and Ta, a trough at K and Pb, and a smooth decreasing slope towards the HREE, with no trough  
248 at Zr and Hf, confirming the similarity of the trace element enrichment in the mantle sources  
249 throughout the Cenozoic province (cf. Melluso et al., 2016). The very marked geochemical  
250 differences with the primitive (MgO=9-12 wt.%) *haiüyne-leucite*-basanites of Mt. Vulture (Italy),  
251 which were generated in a subduction-modified upper mantle source (e.g., Beccaluva et al., 2002),  
252 and the similarities with the potassic basanites of the Virunga area (Congo-Rwanda, western branch  
253 of the East African Rift), another within-plate volcanic area as is Itasy (e.g., Rogers et al., 1992), are  
254 highlighted in Fig. 5.

255

## 256 **5. The concentration of trace elements in titanite, clinopyroxene, kaersutite and anorthoclase** 257 **of Itasy trachyphonolites**

258

259 The mantle normalized patterns of the evolved rocks (Fig. 5c) show relative decrease of P, Ti, Ba,  
260 Sr, Ta, Y and middle-REE, and relative increase of Rb, Th, U, Nb, Zr and Hf from the least to the  
261 most evolved trachyphonolites. In order to understand the changes of the chemical composition of  
262 the trachyphonolites, we also investigated the geochemistry of titanite, kaersutite, clinopyroxene and  
263 anorthoclase phenocrysts of the trachyphonolite RT-06I-355 (Table 2; Fig. 6; Supplementary Fig.  
264 S8). The **titanites** have very high concentration of REE ( $\Sigma REE=28,000-31,500$  ppm), Y (1730-2250  
265 ppm), Ta (650-1130 ppm), Th (134-187 ppm), V (216-328 ppm) and Hf (140-202 ppm), high Zr  
266 (4500-5800 ppm), Nb (6200-15230 ppm), U (15-27 ppm) and La/Yb<sub>n</sub> ratios (19-28), no troughs at  
267 Eu in the chondrite-normalized REE diagrams (Supplementary Fig. S8b), and low Ba (42-49 ppm),  
268 Pb (<1.1 ppm) and Sr (54-382 ppm). The Nb/Ta ratio of the titanites is low (8-13) and the Th/U is  
269 high (8.5-10.3), with respect to typical chondritic ratios (15 and 3.6-4, respectively). The titanites  
270 also have slight chemical zoning, with an increasingly concave REE pattern (Table 2;  
271 Supplementary Fig. S8b). The patterns of the Itasy titanites are generally similar, but displaced to  
272 higher REE concentrations, to those reported by Marks et al. (2008) and Melluso et al. (2010) on  
273 titanites of alkaline intrusive rocks of variable degree of evolution (clinopyroxenites, ijolites,  
274 monzosyenites, to agpaitic nepheline syenites; Supplementary Fig. 8g), which also have widely  
275 variable Nb/Ta ratios (8-66). The Itasy titanite/bulk-rock elemental ratios are of the same magnitude  
276 of the partition coefficients of Olin and Wolff (2012) only for La, Ce and Lu (Supplementary Fig.  
277 8f). **Kaersutite** is a corroded phase, and has a different trace element pattern, with a concentration of

278  $\Sigma$ REE which is one or two orders of magnitude lower than the concentration of REE of the  
279 coexisting titanites (911 ppm;  $La_n = 419$  times chondrite; Table 2). It still has relatively high Li (144  
280 ppm), Y (81 ppm), Zr (732 ppm), Nb (200 ppm), moderate Ba (1100 ppm) and very low Rb, Pb, Th  
281 and U, when compared to the host-rock bulk composition (Table 1; Fig. 6b). The **clinopyroxene** has  
282 low concentration of REE ( $\Sigma$ REE=350-690 ppm;  $La/Yb_n = 8-15$ ), and very low concentration of other  
283 elements (Table 2), excluding Zn (600-662 ppm). **Anorthoclase** has high and variable Ba (1500-  
284 28220 ppm), Sr (534-4300 ppm), Rb (7-98 ppm), Pb (4.6-9 ppm) and Eu (0.5-1.6 ppm). These  
285 elements peak in the mantle-normalized diagrams (Fig. 6b). Other chemical details can be found in  
286 the Table 2.

287

## 288 **6. Isotopic composition of the Itasy rocks**

289

290 The isotopic composition of Itasy rocks is reported in the Table 3. The basanites (MgO= 7-10 wt.%)  
291 have a restricted range of  $^{87}\text{Sr}/^{86}\text{Sr}$  (0.70366-0.70378),  $^{143}\text{Nd}/^{144}\text{Nd}$  (0.51274-0.51277) and Pb  
292 isotopes ( $^{206}\text{Pb}/^{204}\text{Pb}=18.7-18.9$ ;  $^{207}\text{Pb}/^{204}\text{Pb}=15.53-15.56$ ;  $^{208}\text{Pb}/^{204}\text{Pb}=38.89-39.01$ ), which indicate  
293 a largely uniform mantle source, characterized by low time-integrated Rb/Sr and Sm/Nd ratios. This  
294 source is isotopically similar, but not identical, to that of the mafic alkaline rocks of the Ankaratra  
295 (nephelinites, basanites and alkali basalts; cf. Melluso et al., 2016; Cucciniello et al., 2017; Fig. 7)  
296 and to that of the olivine melilitites to the east and northeast of Ankaratra (Melluso et al., 2011a),  
297 which do not have evidence of crustal contamination. The phonotephrites have  $^{87}\text{Sr}/^{86}\text{Sr}=0.70394-$   
298  $0.70398$ ,  $^{143}\text{Nd}/^{144}\text{Nd}= 0.51273-0.51274$ ,  $^{206}\text{Pb}/^{204}\text{Pb}=18.74-18.84$ ,  $^{207}\text{Pb}/^{204}\text{Pb}=15.53-15.54$  and  
299  $^{208}\text{Pb}/^{204}\text{Pb}=38.97-39.02$ , all values broadly similar (or indistinguishable in some isotope ratios) to  
300 those of the basanites, whereas the benmoreite and the trachyphonolites have higher  $^{87}\text{Sr}/^{86}\text{Sr}$   
301 (0.70430-0.70446) and lower  $^{143}\text{Nd}/^{144}\text{Nd}$  (0.51266-0.512682),  $^{206}\text{Pb}/^{204}\text{Pb}$  (18.18-18.39),  
302  $^{207}\text{Pb}/^{204}\text{Pb}$  (15.49-15.51) and  $^{208}\text{Pb}/^{204}\text{Pb}$  (38.38-38.57) than the mafic rocks. Therefore, the isotopic  
303 composition of the Itasy rocks changes with MgO. A broadly similar isotopic variation between  
304 associated mafic and evolved rocks has been noted in both Cretaceous and Cenozoic volcanic rocks  
305 of Madagascar (cf. Cucciniello et al., 2010, 2013, 2017) and indicates that, if basanites and  
306 trachyphonolites (which crop out in the same area and have the same stratigraphic age) are part of a  
307 comagmatic suite, the evolved rocks suffered interaction with the old, low-Pb/Pb basement rocks  
308 during magmatic evolution in low-pressure magma reservoirs. This part is developed in the  
309 discussion.

310

## 311 **7. Discussion**

312

### 313 7.1. The unusual transition between basanites/tephrites and trachyphonolites

314

315 The Itasy volcanic rocks have some interesting petrogenetic features shared with those of other  
316 volcanic areas worldwide: 1) the leucite-bearing basanites recall analogues found in other areas (e.g.,  
317 the Roman Province, the western branch of the East African Rift); 2) the presence of haüyne  
318 ( $\pm$ sodalite)-bearing alkaline volcanic rocks is found in alkaline rocks of various tectonic settings  
319 (e.g., Massif Central: Vatin-Perignon, 1968; Eifel: Wörner and Schmincke, 1984; Tenerife: Bryan,  
320 2006; Mt. Etinde: Nkombou et al., 1995; Mt. Vulture: Beccaluva et al., 2002; Melluso et al., 1996,  
321 2011b, among other places), and indicate abundant sulfur in the oxidized state (and chlorine as  
322 well); 3) trachyphonolites similar to those of Itasy are common evolved lithotype of other major  
323 volcanic areas (e.g., Suswa, Kenya; Phlegrean Fields-Ischia, Italy). These trachyphonolites are the  
324 product of low-pressure fractional crystallization processes, possibly in an open system. From a  
325 more general view, basanites and tephrites (and maybe also some nephelinite types) are associated  
326 with phonolites *s.s.*, rather than trachyphonolites, as is found in northernmost Madagascar and  
327 elsewhere (e.g., le Roex et al., 1990; Ablay et al., 1998; Melluso et al., 2007; Berger et al., 2014). In  
328 contrast, trachytes (trachyphonolites) are typically associated with (potassic or sodic) alkali and  
329 transitional basalts (e.g., Marion Island, le Roex et al., 2012; St. Helena, Kawabata et al., 2010 and  
330 references therein; Ischia-Phlegrean Fields, Fedele et al., 2008; Melluso et al., 2012, 2014; Libya,  
331 Lustrino et al., 2012; Mauritius: Sheth et al., 2003; Ashwal et al., 2016; Ethiopian Rift, Ronga et al.,  
332 2010; Kenya Rift: White et al., 2012).

333 The magmatic evolution of the Itasy basanites and tephrites (the most Mg-rich of them equilibrated  
334 with olivine at temperatures just lower than 1200°C, according to Roeder and Emslie, 1970), is  
335 without any evidence of removal of plagioclase or leucite, because these phases are found  
336 exclusively as tiny groundmass microlites (e.g., Supplementary Fig. 1). Sodic plagioclase appears as  
337 a phenocryst phase in benmoreites; anorthoclase and Na-sanidine (plus rare mafic phases) are the  
338 dominant phenocrysts and groundmass phases in the trachyphonolites. It is thus evident that removal  
339 of Na-feldspars was essential for the evolution of (sodic) trachyphonolitic liquids towards the  
340 phonolite minimum. This is also indicated by the position of the evolved samples in the Petrogeny  
341 Residua's System, which plot between the trachytic and the phonolitic minima (slightly less than  
342 900°C, at 1 kbar  $P_{H_2O}$ ), hence within the silica undersaturated field (Fig. 8a, b). The tendency of the  
343 bulk-rock trachyphonolitic compositions to plot close to the composition of their (Na-rich) alkali  
344 feldspars (Fig. 4a) can be considered as the *normal* evolution towards the pertinent minimum melt  
345 compositions. The lack of *phenocrysts* of sodic feldspathoids is another indication that the bulk-rock  
346 compositions approach, but do not reach, the phonolitic minimum. For comparison, the phonolites

347 (s.s.) of northernmost Madagascar (Melluso et al., 2007; Cucciniello et al., 2016) have nepheline  
348 phenocrysts, and plot on the alkali feldspar-nepheline cotectics (Fig. 8b).  
349 The volatiles that left trace in the Itasy magmatic system are H<sub>2</sub>O (hosted in amphibole, apatite, the  
350 rare biotite, and glass), SO<sub>3</sub> (hosted in the haüyne-sodalite solid solutions and in apatite), Cl (hosted  
351 in sodalite, apatite and glass) and F (hosted in apatite, amphibole, glass, fluorite and other  
352 accessories). No evidence for a significant role of CO<sub>2</sub> in the petrogenesis of the Itasy basanites and  
353 trachyphonolites has been found, as expected by the insignificant maximum quantity of this  
354 compound that can be theoretically dissolved in basanitic melts at low pressure, when compared to  
355 H<sub>2</sub>O (e.g., Shishkina et al., 2014), a quantity even lower in trachytes and phonolites (e.g., Webster et  
356 al., 2014). The Itasy basanites and tephrites have significant amounts of kaersutite, indicating that  
357 the magmatic crystallization took place in a H<sub>2</sub>O-rich environment at relatively early differentiation  
358 stages. Kaersutite is unknown in basanites and alkali basalts (and nephelinites too) of the nearby  
359 Ankaratra complex, that have mostly anhydrous mineral assemblages, and occurs as a phenocryst  
360 phase only in a few benmoreites (Cucciniello et al., 2017). Kaersutite s.s. crystallizes in mafic,  
361 hydrous, feldspar-bearing (or -normative), high-Ti alkaline magmas, and has a range of  
362 crystallization depths from the uppermost mantle (e.g., Irving and Green, 2008; Pilet et al., 2010) to  
363 shallow conduits or shallow intrusions (e.g., Melluso et al., 2005, 2007); hence this phase is more a  
364 chemical than a barometric indicator (cf. Supplementary Fig. S4). Sulfur was at least partially  
365 partitioned in the S-bearing apatite before reaching saturation as haüyne, in the groundmass of the  
366 phonotephrites and in the evolved rocks; sulfur decreases in the latest crystallized groundmass, as  
367 testified by the presence of sodalite (cf. also Supplementary Fig. S1).

368

## 369 **7.2. Petrogenetic modelling: the essential role of titanite and anorthoclase removal in the** 370 **geochemical evolution of the trachyphonolites, and open-system processes**

371

372 The petrogenetic study of the Itasy volcanic rocks is focused on the genetic relationships between  
373 this bimodal suite of mafic and evolved rocks, with their changing isotopic ratios and the marked  
374 decrease in the degree of silica undersaturation of the evolved rocks. Despite the compositional gap  
375 between mafic and evolved rocks shown in the diagrams, involving a marked change of density,  
376 temperatures and time of storage in shallow reservoirs, we observe the following: 1) the mineral  
377 phases have a continuous compositional change from mafic, through intermediate, to felsic rocks, as  
378 a typical consequence of fractional crystallization processes (see above); 2) the bulk-rock  
379 compositions have the expected changes considering the variety and amount of the phenocryst  
380 phases, thus excluding effects of magma mixing; 3) many basanites and tephrites are rich in  
381 phenocrysts/megacrysts of kaersutite (a phase significantly lower in silica than clinopyroxene but



382 with similar density), whose significant removal could help increase the concentration of SiO<sub>2</sub> of  
383 residual melts.

384 The variation diagrams (Supplementary Fig. S6) and the petrography of the mafic rocks indicate that  
385 clinopyroxene, amphibole, olivine and magnetite (for Fe and Ti) are probably the best candidates to  
386 form the subtracted mineral assemblages in the transition from basanites to phonotephrites. Among  
387 several models using the major element compositions of bulk-rocks and their phenocryst phases, the  
388 transition from basanites (e.g. RT-06I-354B) to tephrites (RT-06I-395) and benmoreite (RT-06I-  
389 387) can be accounted for by 37% and 56% removal of ultramafic assemblages, respectively. The  
390 transition from the benmoreite RT-06I-387 to the trachyphonolite RT-06I-375B needs 61% removal  
391 of monzonitic assemblages. The total transition from the basanite RT-06I-354B to a trachyphonolite  
392 (e.g., sample RT-06I-375B) was modelled with 71.4% removal of ultramafic assemblages (cf.  
393 Supplementary Table S10). The results we choose have low  $\Sigma R^2$  (values of  $\Sigma R^2 \ll$  or  $<0.5$  can be  
394 considered for modelling, but they are not proof of geological reliability), but the high  
395 kaersutite/clinopyroxene ratio involved in some transitions between mafic rocks is always to be  
396 matched with the actual subtracted assemblages, which are, of course, largely unknown.

397 Thermodynamic models of magmatic evolution (e.g., MELTS) are not able to account for the  
398 presence of kaersutite, sodalite-group minerals, titanite and a number of other petrographic features  
399 of alkaline rock suites, including the crystallization sequence, the phase composition and their actual  
400 amount. LeMasurier et al. (2011) hypothesized a significant role for kaersutite removal in order to  
401 account for the petrogenesis of evolved silica oversaturated rocks starting from basanite parental  
402 magmas; in our case, we show that this effect must be limited to the transition from tephrites to  
403 benmoreites, because the final transition to the most evolved Itasy trachyphonolites cannot involve  
404 significant removal of amphibole (kaersutite is already a minor phenocryst phase in the least evolved  
405 trachyphonolites).

406 A number of observations and models indicate that the Itasy benmoreites and trachyphonolites could  
407 be comagmatic with the basanites, through (also) the effects of kaersutite removal. The trace element  
408 variations using thorium as the most incompatible element were modelled using the Rayleigh  
409 equation (Supplementary Fig. S7). The following features are of interest: a) assuming that Th  
410 behaves in a closed system (an approximation that cannot be guaranteed during AFC processes, see  
411 below) and that is the most incompatible element, the total amount of crystallization needed to reach  
412 the phonotephritic liquids is roughly 60%, and is 75-80% to reach evolved trachyphonolites, starting  
413 from the least evolved basanites; b) the transition from basic to intermediate rocks is characterized  
414 by several major changes in the slope of the trace elements vs. the Th concentration and, therefore,  
415 in their bulk partition coefficients (D); this change in slope takes place at different evolution points  
416 (Supplementary Fig. S6 and S7); c) tantalum (Ta) behaves as a highly compatible element in the

417 trachyphonolites, thus strongly decouples from the geochemically similar element Nb; roughly  
418 similar fractionation is noted between Hf and Zr, whereas U and Th behave sympathetically (the  
419 Th/U ratio ranges between 3.6 and 4.4, both values in basanites); d) lead (Pb) increases more than  
420 that expected from a perfectly incompatible behaviour (modelling requires  $D_{Pb} < 0$ ; Supplementary  
421 Fig. S7), strongly suggesting the introduction of external Pb during crustal contamination processes;  
422 e) the chemical composition of the trachyphonolites is scattered, indicating that the magmatic  
423 evolution took place in independent magma feeder systems, hence with variable distribution,  
424 crystallization and removal of phases.

425 The petrography, the increasingly concave REE patterns of the trachyphonolites, the trace element  
426 concentration of the titanite phenocrysts, and numerical modelling using both published partition  
427 coefficients and the mineral composition determined by LA-ICP-MS (Supplementary Fig. S8)  
428 indicate a major effect of titanite on the fractionation of middle-REE, Y, Ta and other elements,  
429 even though full quantification of the extent of removal of an accessory phase cannot be obtained  
430 through major oxide mass balance calculations or trace element behaviour. The trend to the most  
431 evolved trachyphonolites can be obtained through 0.7-1% removal of titanite (cf. Supplementary  
432 Fig. S8c, d). Removal of small amounts of titanite is commonplace in evolved alkaline rocks (cf. the  
433 extremely similar patterns reported in Ulrych et al., 2006; Lustrino et al., 2012; Melluso et al., 2014,  
434 2017 and references therein), and the effects observed in this work are broadly those expected  
435 utilizing known titanite/melt partitioning data (e.g., Olin and Wolff, 2012; Supplementary Fig. S8c,  
436 Supplementary Table S11). Apatite can have a roughly similar behaviour regarding middle-REE  
437 fractionation, although to a lower extent (the partition coefficients of apatite can be lower of a factor  
438 of 3-4; cf. Supplementary Table S11 and Supplementary Fig. S8c), but has negligible effect on the  
439 concentration of Ta, Hf, Zr and Nb. Removal of other phases such as clinopyroxene or amphibole, in  
440 the amount observed in the least evolved trachyphonolites (Supplementary Fig. S1), cannot have  
441 appreciable role in preferentially removing middle- from heavy-REE in evolved melts (cf.  
442 partitioning data of Tiepolo et al. 2007; Fedele et al., 2009; Melluso et al., 2017; this work), and so  
443 is for other accessory phases *absent* at Itasy (e.g., melanite garnet, or significant zircon).

444 The evolved samples of Itasy show compelling evidence for open-system evolution, as seen from the  
445 isotopic variations and elemental evidence above. We performed an AFC model (DePaolo, 1981)  
446 utilizing all the isotopic systematics, pertinent trace elements and compositions of the Madagascan  
447 basement in the area. The results highlight the effects of prolonged fractional crystallization (roughly  
448 70%) and limited, though widely variable, crustal contamination (the values of the  
449 assimilation/cumulus rate,  $-r$ , are  $< 0.3$ ) of known felsic rocks of the Precambrian basement (details  
450 in Fig. 9). The model takes into account the marked changes in the bulk partition coefficients of  
451 many elements at varying levels of magmatic evolution (Supplementary Fig. S8). The

452 phonotephrites (samples RT-06I-397 and RT-06I-398) still need limited amount of crustal  
453 contamination in order to account for their slightly differing isotopic ratios compared to those of  
454 basanites and tephrites. The abrupt change of partition coefficients in the evolved rocks and a limited  
455 change of the parameter “r” (Fig. 9) led to the bimodal isotopic compositions. It is also evident that  
456 the effects of crustal contamination did not act uniformly throughout the Itasy suite, and that other  
457 processes could have been operative in the Itasy feeder system, including silica increase that may be  
458 expected from interaction with the felsic lithologies of the Madagascan basement, and the presence  
459 of alkali basaltic parental magmas (more SiO<sub>2</sub>-rich than a basanite), of which there is no evidence, *to*  
460 *date*.

461

### 462 **7.3. The mantle sources of Itasy in the context of the Cenozoic magmatism of Madagascar**

463

464 The Itasy basanites have Sr-Nd-Pb-isotopic ratios indicating a source geochemically and isotopically  
465 different from that of convecting asthenosphere, which experienced a distinct evolution with time.  
466 This is fully consistent with what is known in the Cenozoic mafic rocks of the central Madagascar  
467 (Ankaratra, Takarindiona and Alaotra), which were emplaced through Archean crust (Melluso et al.,  
468 2011a, 2016; Cucciniello et al., 2017). Primitive basanitic magmas with the petrographic and  
469 geochemical characteristics similar to those of the Itasy imply low degrees of partial melting (ca.  
470 3%) of a hydrous, amphibole-bearing lherzolite, and hence at depths and temperatures where  
471 *pargasitic* amphibole can be a stable phase of a peridotite (say, <90-100 km; e.g., Kovacs et al.,  
472 2017), as previously suggested by earlier work in central and northern Madagascar (Melluso and  
473 Morra, 2000; Melluso et al., 2016 and references therein; Rocco et al., 2017), and elsewhere (cf. le  
474 Roex et al., 2001; Green and Falloon, 2015). A relatively shallow depth of the (lithospheric) source  
475 region of the Itasy basanites is also indicated by a rootless low-velocity zone (supposedly melt-,  
476 volatile- or light phases-bearing) just beneath the Moho, extending at depths of 50-80 km (Fig. 1c;  
477 Pratt et al., 2017). The E-W-trending extensional processes active in central Madagascar seem to  
478 have favoured moderate uplift of this geochemically enriched, volatile-bearing lithospheric mantle,  
479 bringing it above the solidus and forming the Itasy low-volume basanitic lava field. The source of  
480 the Itasy basanites must be characterized by a significant enrichment of highly incompatible  
481 elements (cf. the low Ba/Nb and Zr/Nb ratios, and the shape of the mantle normalized patterns), with  
482 particular reference to the HFSE, in agreement with a fully within-plate origin, without subduction-  
483 related components or low-pressure crustal contamination. The source of the Itasy basanites cannot  
484 be depleted asthenospheric -or primitive- mantle, assuming any reasonable degrees of partial melting  
485 and considering the isotopic composition (cf. also Melluso et al., 2016 and Cucciniello et al., 2018,  
486 for a regional review). Finally, in terms of trace element and isotopic composition, the sources of the

487 Cenozoic volcanic rocks of Madagascar share very little with those giving rise the Cretaceous  
488 *tholeiitic* magmatism, implying different melting regimes of differently located mantle sources (cf.  
489 Melluso et al., 2001, 2005, 2011a, 2016; Cucciniello et al., 2010, 2013, 2017).

490

## 491 **8. Conclusions**

492

493 The Itasy volcanic field, the most recently active in Madagascar, is characterized by an association  
494 of high-TiO<sub>2</sub>, potassic, basanites/tephrites/phonotephrites, and benmoreites/ trachyphonolites, in the  
495 facies of pyroclastic rocks and lava flows, with the highly featuring presence of Ti-rich amphibole  
496 (and leucite) in mafic/intermediate rocks and haüyne-sodalite in evolved rocks. Recognizing the  
497 possibility of independent basanitic parental magmas and different liquid lines of descent, leucite  
498 basanite magmas evolved through removal of olivine, clinopyroxene, kaersutite and magnetite to  
499 reach the volumetrically minor benmoreitic and trachyphonolitic compositions. These latter rocks  
500 are likely the result of low-pressure fractional crystallization of anorthoclase/sanidine-bearing  
501 (*syenitic*) assemblages and minor crustal contamination of the high-<sup>87</sup>Sr/<sup>86</sup>Sr, low-<sup>143</sup>Nd/<sup>144</sup>Nd, -  
502 <sup>206</sup>Pb/<sup>204</sup>Pb, -<sup>207</sup>Pb/<sup>204</sup>Pb and -<sup>208</sup>Pb/<sup>204</sup>Pb, old granitic-migmatitic basement of the area, with the  
503 evidence of minor titanite removal (perhaps <2%), as part of syenitic/monzosyenitic cumulates.  
504 Titanite is demonstrated to be a very efficient scavenge of intermediate-REE, Y, Ta and V, and  
505 slightly less so for Nb, Zr and Hf in the Itasy trachytic (*s.l.*) magmas (cf. Table 2); its removal, even  
506 in small amounts, significantly increases ratios between otherwise geochemically similar elements  
507 (e.g., Zr/Hf, Nb/Ta) when compared to those of the associated mafic rocks. Another minor but  
508 interesting result of this study is the evidence that the concentration of rare earths or other strategic  
509 elements in the most highly evolved alkaline magmas cannot be considered a rule working  
510 worldwide, due to the presence of accessory phases on the liquidus, which effectively scavenge  
511 these elements in less evolved melts. *Also* due to the removal of titanite and other accessory phases,  
512 the geochemical characteristics of trachytic (*s.l.*) rocks *worldwide* cannot give unambiguous  
513 indication about those of associated (or parental) mafic magmas, or even to indicate mantle  
514 signatures of any kind, including ratios of elements thought to be highly incompatible. The  
515 geochemical and isotopic characteristics of the leucite-basanites broadly match those of the nearby  
516 Ankaratra volcano and of the rest of the Cenozoic volcanic districts of Madagascar (Melluso et al.,  
517 2016), and testify to a mantle source having marked enrichment in LILE, HFSE and volatile  
518 elements (H<sub>2</sub>O, Cl, F, SO<sub>3</sub>), as noted in other within-plate strongly alkaline rocks worldwide and,  
519 intriguingly, also in some subduction-related alkaline rocks. The broadly E-W extensional processes  
520 common in central Madagascar triggered low-degree partial melting of enriched, volatile-rich,  
521 lithospheric mantle, and generated the low-volume Itasy basanites, without the need to invoke the

522 action of mantle plumes. Enriched geochemical components of within-plate origin are commonplace  
523 in the Madagascan lithospheric mantle (Melluso et al., 2016; Rocco et al., 2017). Finally, the  
524 unusual petrogenetic association between basanites and trachyphonolites, and the likely significant  
525 removal of kaersutitic amphibole in some basanites and tephrites, are of importance in identifying  
526 the volatile species of within-plate alkaline rocks in the mantle sources and during the magmatic  
527 evolution.

528 A high-resolution geochronological framework is now needed to establish an accurate age span for  
529 the Itasy, if a significant temporal lapse exists between mafic and evolved rocks, and the  
530 relationships between timing of volcanism and extensional tectonics of central Madagascar.

531

### 532 **Acknowledgements**

533

534 This paper is dedicated to the memory of Fabio Carmine Mazzeo.

535 Alessio Langella, Sergio Bravi, Petrus Le Roux, Roberto de' Gennaro, Vohangy Ratrimo,  
536 Dieudonné Razafimahatratra, Vincenza Guarino, Luigi Franciosi and Lorenzo Fedele are gratefully  
537 thanked for helping us in sampling, data acquisition and comments on an early draft of the  
538 manuscript. The constructive reviews of Jean-Paul Liégeois and Michael Marks and advice of the  
539 Editor-in-chief Xian-Hua Li were very helpful to improve contents and presentation of an early  
540 version of the manuscript.

541 This paper was supported by grants provided by PRIN2015 (grant 20158A9CBM) and LR5  
542 (Regione Campania) to L. Melluso, and Fondi Ricerca di Ateneo (DR\_3450\_2016) to C.  
543 Cucciniello.

544

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782

### 783 **Appendix: analytical techniques**

784

785 Powders of the 69 samples of this study were obtained after carefully grinding washed chips in an  
786 ultrapure agate mill. For each sample, four grams of micronized powder were used to prepare pressed  
787 powder pellets. The powder (mixed with 1 ml of Polyvinyl alcohol solution) was pressed to twenty  
788 tons for twenty seconds. The bulk-rock X-Ray Fluorescence compositional data (Table S1) were  
789 obtained on pressed powder pellets with an Axios Panalytical spectrometer equipped with six analyzer  
790 crystals, three primary collimators and two detectors (flow counter and scintillator), operating at  
791 different kV and mA for each analyte. Analytical uncertainties are in the order of 1-2% for major  
792 elements and 5-10% for trace elements. The weight loss on ignition has been obtained with  
793 gravimetric techniques, firing at 1000°C small aliquots of powders previously dried at 110°C  
794 overnight.

795 Data on a subset of samples were obtained through ICP-MS methods at Actlabs (Canada) (Table 1).  
796 Samples were mixed with a flux of lithium metaborate and lithium tetraborate and fused in an  
797 induction furnace. The molten material is immediately poured into a solution of 5% nitric acid  
798 containing an internal standard, and mixed continuously until completely dissolved (~30 minutes). The

799 samples were run for major oxides and selected trace elements (Ba, Be, Sc, Sr, V, Y and Zr) on a  
800 combination simultaneous/sequential Thermo Jarrell-Ash ENVIRO II ICP or a Varian Vista 735 ICP.  
801 Calibration is performed using seven prepared USGS and CANMET certified reference materials. One  
802 of the seven standards is used during the analysis for every group of ten samples. Sample fused are  
803 diluted and analyzed by Perkin Elmer Sciex ELAN 6000, 6100 or 9000 ICP/MS for other trace  
804 elements (Cr, Co, Ni, Cu, Zn, Ga, Ge, As, Rb, Nb, Mo, Ag, In, Sn, Sb, Cs, La, Ce, Pr, Nd, Sm, Eu,  
805 Gd, Tb, Dy, Ho, Er, Tm, Yb, Lu, Hf, Ta, W, Tl, Pb, Bi, Th and U). Three blanks and five controls  
806 (three before the sample group and two after) are analyzed per group of samples. Duplicates are fused  
807 and analyzed every 15 samples. Analyses of international standards are reported in Cucciniello et al.  
808 (2017).

809 Microprobe analyses were carried out on polished thin sections using Energy Dispersive Spectrometry  
810 (EDS) at University of Napoli Federico II, utilizing a JEOL JSM-5310 microscope operating at 15 kV  
811 primary beam voltage, 50-100 mA filament current, 50 s net acquisition time and a Oxford  
812 Instruments Microanalysis Unit, equipped with an INCA X-act detector. For further details see  
813 Melluso et al. (2014, 2017).

814 The Laser Ablation - Inductively Coupled Plasma - Mass Spectrometry analysis of the minerals in the  
815 trachyphonolite RT-06I-355 was performed at IGG-CNR (Pavia, Italy), utilizing techniques described  
816 in Rocco et al. (2017). Concentrations of Rare Earth Elements (REE) and other selected trace elements  
817 (Li, B, Ba, Rb, Th, U, Nb, Ta, Sr, Zr, Hf, Ti, Y, Sc, V, Co and Ni) were determined in situ on polished  
818 sections (100  $\mu\text{m}$ ) using a PerkinElmerSCIEX ELAN DRC-e quadrupole mass spectrometer coupled  
819 with an UP213 deep-UV YAG Laser Ablation System (New Wave Research, Inc.). The laser was  
820 operated at a repetition rate of 10 Hz, with 213 nm wavelength and a fluence of  $\sim 9.5 \text{ J/cm}^2$ . Helium  
821 was used as carrier gas and was mixed with Ar downstream of the ablation cell. Spot diameter was 55  
822  $\mu\text{m}$ . Data reduction was performed offline using the GLITTER software. For this study, the NIST  
823 SRM 610 synthetic glass standards was used as external standard, CaO was used as internal standard  
824 for clinopyroxene and amphibole, while  $\text{SiO}_2$  was used for titanite and feldspar. Precision and  
825 accuracy of the REE concentration values were assessed through repeated analysis of the BCR2-g  
826 standard to be better than  $\pm 7\%$  and  $\pm 10\%$ , respectively, at the ppm concentration level (for further  
827 details see Miller et al., 2012 and references therein).

828 The Sr-Nd-Pb-isotope data were obtained at the University of Cape Town, with techniques described  
829 in le Roex et al. (2012). Sr, Nd and Pb were separated using conventional ion exchange techniques and  
830 all radiogenic isotope analyses were performed on a NuPlasma multicollector inductively coupled  
831 plasma-mass spectrometer (ICP-MS). To correct for mass fractionation effects, measured  $^{87}\text{Sr}/^{86}\text{Sr}$  and  
832  $^{143}\text{Nd}/^{144}\text{Nd}$  values were normalized to  $^{86}\text{Sr}/^{88}\text{Sr}=0.1194$  and  $^{146}\text{Nd}/^{144}\text{Nd}=0.7219$ , respectively. Pb

833 isotopes were corrected for fractionation by normalizing ratios measured in international standards.  
834 Average standard values obtained during the course of this study are reported in the Table 3.

835

### 836 **Table captions**

837

838 Table 1: ICP-MS major and trace element composition of Itasy rocks. A few element ratios are also  
839 reported.

840 Table 2: LAM-ICP-MS and microprobe composition of titanite, amphibole, clinopyroxene and  
841 feldspar of the trachyphonolite RT-06I-355. The location of the laser pits is shown in the  
842 Supplementary Fig. S1c.

843 Table 3: Isotopic composition of the Itasy rocks. The composition of the standards which were run  
844 with the unknowns is also reported.

845

### 846 **Supplementary Tables**

847

848 Supplementary Table S1: XRF major trace element data and CIPW norms of the Itasy samples. The  
849 composition of international standards is also reported.

850 Supplementary Table S2: synopsis of the mineral assemblages of the main lithotypes of the Itasy  
851 complex.

852 Supplementary Table S3: composition of olivine of the Itasy rocks.

853 Supplementary Table S4: composition of oxides of the Itasy rocks.

854 Supplementary Table S5: composition of pyroxene of the Itasy rocks.

855 Supplementary Table S6: composition of amphibole, biotite and rhönite of the Itasy rocks.

856 Supplementary Table S7: composition of feldspar and glass of the Itasy rocks.

857 Supplementary Table S8: composition of feldspathoids of the Itasy rocks.

858 Supplementary Table S9: composition of titanite, apatite, other accessories of the Itasy rocks.

859 Supplementary Table S10: Recapitulation of mass balance calculations between rocks of different  
860 degree of magmatic evolution, and detailed results. The composition of the phases is reported in  
861 the supplementary tables.

862 Supplementary Table S11: average REE mineral/bulk rock ratios for titanite, amphibole and  
863 clinopyroxene. The partition coefficients of titanite, clinopyroxene, apatite and amphibole are  
864 taken from Olin and Wolff (2012), Fedele et al. (2009, 2015) and Tiepolo et al. (2007). Note that  
865 amphibole of RT-06I-355 could not be considered in equilibrium with the host rock (cf.  
866 Supplementary Fig. S1).

867

868 **Figure captions**

869

870 Figure 1: a) Simplified geological map of Madagascar, with the location of the main volcanic areas  
871 of Cretaceous and Cenozoic; b) Geological sketch map with sample localities (Laborde  
872 coordinates) and the main tectonic features of the Itasy area. c) Schematic section of the Itasy  
873 region (redrawn after Pratt et al., 2017) with location of a possible mantle source region of the  
874 Itasy mafic volcanic rocks.

875 Figure 2: Classification of the Itasy rocks according to the TAS and  $R_1R_2$  diagrams. Symbols: Itasy  
876 mafic rocks: red triangles; Itasy evolved rocks: blue circles. The data from other Cenozoic  
877 districts of Madagascar are taken from: Melluso et al. (2007, 2011b); Cucciniello et al. (2011,  
878 2016, 2017); Melluso and Morra (2000).

879 Figure 3: Classification of pyroxene and amphibole of the Itasy rocks. Symbols as in Fig. 2. The  
880 pyroxene and amphibole composition of the Ankaratra alkaline rocks (black squares) are taken  
881 from Cucciniello et al. (2017).

882 Figure 4: a) The composition of feldspars and bulk-rocks of the Itasy in the An-Ab-Or diagram  
883 (mol.%). b) Composition of the Itasy volatile-rich feldspathoids in the CaO-Na<sub>2</sub>O-K<sub>2</sub>O diagram  
884 (mol.%; after Lessing and Grout, 1971). c) The volatile concentration of the Itasy feldspathoids  
885 (in wt.%). The feldspar composition of the Ankaratra alkaline rocks (black squares) are taken  
886 from Cucciniello et al. (2017). The data on Mt. Vulture feldspathoids are taken from Melluso et  
887 al. (1996, 2011a), references therein and unpublished data. Symbols as in Fig. 2.

888 Figure 5: a, b) Mantle normalized diagrams for the Itasy basanites. The data on the other  
889 Madagascan basanites (grey patterns in the background) are taken from the database of Melluso  
890 et al. (2016). The analyses of the Mt. Vulture basanites (MgO=9-12 wt.%) are taken from  
891 Beccaluva et al. (2002), L.M. and V.M., unpublished data. The data on the Karisimbi leucite  
892 basanites (green patterns in the background) are taken from Rogers et al. (1992). At a similar  
893 MgO, the mantle-normalized patterns of the Mt. Vulture and Itasy basanites (as well other  
894 geochemical characteristics) imply contrasting enrichment processes in their respective sources,  
895 and a different mantle baseline preceding metasomatic events; c) Mantle normalized diagrams  
896 for the Itasy evolved rocks. The normalization values are taken from Lyubetskaya and Korenaga  
897 (2007).

898 Figure 6: a, b) Mantle normalized diagrams of the coexisting phases of the trachyphonolite RT-06I-  
899 355, together with those of the Itasy evolved rocks. Normalization values are those of  
900 Lyubetskaya and Korenaga (2007). Potassium and Phosphorus are not reported.

901 Figure 7: Isotopic diagrams for the Itasy and the other Cenozoic volcanic rocks emplaced within the  
902 cratonic domains of the Madagascan basement (cf. Fig. 1a). Symbols as in Fig. 2. The Ankaratra

903 data are taken from Cucciniello et al. (2017), and the Alaotra and Takarindiona olivine melilitites  
904 (black triangles) are taken from Melluso et al. (2011a).

905 Figure 8: a) The tetrahedral phase diagram diopside-nepheline-kalsilite-silica at one atmosphere  
906 (after Cucciniello, 2016) with the Itasy bulk-rock compositions (symbols as in Fig. 2); b) The  
907 Petrogeny Residua's System at 1 kbar  $P_{H_2O}$  with the Itasy trachyphonolites (blue circles). The  
908 boundary lines are taken from Gupta (2015). The evolved rocks of the Ankaratra (black squares)  
909 are taken from Cucciniello et al. (2017); the Bobaomby phonolites (black triangles) are taken  
910 from Melluso et al. (2007); the trachytes of Mauritius (grey rhombs) are taken from Ashwal et al.  
911 (2016). Other evolved rocks can be found in Cucciniello et al. (2017). The compositions of glass  
912 of the evolved rocks is reported as light blue circles.

913 Figure 9: AFC modelling after DePaolo (1981). The composition of the contaminants (granitic rocks  
914 and felsic gneisses of the basement around Itasy: data from Kröner et al., 2000; Melluso et al.,  
915 2001; Cucciniello et al., 2017) is the following:  $^{87}Sr/^{86}Sr=0.7113$ ,  $^{143}Nd/^{144}Nd=0.5113$ ,  
916  $^{206}Pb/^{204}Pb=16.62$ ,  $Pb=10$  ppm;  $Sr=499$  ppm;  $Nd=16$  ppm;  $Th=12$  ppm. The bulk partition  
917 coefficients are reported along the curves (cf. Supplementary Fig. S7), together with the values of  
918 the residual liquid fraction (f) and the assimilation/accumulus rate (r). The AFC models in the  
919 diagram  $^{143}Nd/^{144}Nd$  vs.  $Nd$  were performed taking into account the geochemical behaviour of Nd  
920 shown in Supplementary Fig. S7, and two starting compositions (the basanite RT-06I-354B and  
921 the benmoreite RT-06I-387). Symbols as in Fig. 2.

922

## 923 **Supplementary Figures**

924

925 Supplementary Fig. S1a: Petrographic, BSE (backscattered electron images), and field images of the  
926 Itasy rocks: a) basanite RT-06I-354B: phenocrysts of olivine, with Cr-spinels, set in a scoriaceous  
927 mesostasis; b) basanite RT-06I-354B: large phenocrysts of kaersutite enclosing diopsidic  
928 clinopyroxene, parallel nicols; c) phonotephrite RT-06I-398: rare phenocrysts of clinopyroxene  
929 and apatite set in a very fine-grained microlitic mesostasis, parallel nicols; d) basanite RT-06I-  
930 354B: large phenocryst of diopside-to-titanaugite clinopyroxene, parallel nicols; e)  
931 trachyphonolite RT-06I-355; phenocrysts of a zoned anorthoclase and titanite in a fine-grained  
932 feldspar-rich mesostasis, parallel nicols; f) trachyphonolite RT-06I-375: phenocrysts of  
933 (pleochroic) titanite and alkali feldspar in a fine-grained groundmass, parallel nicols. The  
934 horizontal scale of the thin section photomicrographs is 2.5 mm.

935 Supplementary Fig. S1b: g) groundmass of basanite RT-06I-354B: microlites of feldspar,  
936 feldspathoids, unmixed Fe-Ti oxides and clinopyroxene, BSE; h) trachyphonolite RT-06I-361:  
937 phenocrysts of amphibole, and microphenocrysts of titanite and apatite in a fine-grained



938 mesostasis; i) trachyphonolite RT-06I-363: microphenocrysts of Fe-Ti oxides, titanite, and  
939 clinopyroxene in a fine-grained mesostasis rich in alkali feldspar and light grey glass, BSE; j)  
940 phonotephrite RT-06I-398: a microlite of h aüyne (darker grey) in the center of the groundmass  
941 made up of feldspar, clinopyroxene, magnetite and glass, BSE; k) l) m) n) o) feldspathoids in the  
942 groundmass of trachyphonolites RT-06I-361 and RT-06I-379; p), q), r) pyroclastic rocks, lava  
943 domes, and spatter cones in the northeastern part of Itasy.

944 Supplementary Fig. S1c: Petrographic characteristics of the trachyphonolite RT-06I-355. a) zoned  
945 green cpx phenocryst b), c) phenocrysts of alkali feldspar and titanite; d) microlites of alkali  
946 feldspar and altered amphibole. The laser pits after LAM-ICP-MS work (Table 2) can be easily  
947 seen. The horizontal scale of the thin section photomicrographs is 2.5 mm.

948 Supplementary Fig. S2: Composition of the titaniferous magnetite of the Itasy rocks in the MgO-  
949 MnO and TiO<sub>2</sub>-Al<sub>2</sub>O<sub>3</sub> diagrams (in wt.%). Symbols as in Fig. 2.

950 Supplementary Fig. S3: Clinopyroxene composition. Symbols: mafic rocks: red triangles;  
951 benmoreites and trachyphonolites: blue circles. The composition of olivine is plotted in the Ca-  
952 Mg-Fe diagram.

953 Supplementary Fig. S4: Amphibole composition (cations in apfu) indicating the *absence of evidence*  
954 for a high-pressure crystallization environment for this phase, as the calculated Al<sup>VI</sup> is roughly  
955 1:10 the calculated Al<sup>IV</sup> in both mafic and evolved rocks. Symbols as in Fig. 2

956 Supplementary Fig. S5: composition of nepheline in the nepheline-kalsilite-silica diagram. The  
957 nephelines of the olivine melilitites and of the Ankaratra rocks are reported for comparison (data  
958 from Melluso et al., 2011a and Cucciniello et al., 2017).

959 Supplementary Fig. S6: XRF major oxides and trace elements vs. MgO wt.% (plotted on a  
960 logarithmic scale). The composition of olivine, amphibole and clinopyroxene of the basanites  
961 and tephrites is reported in a few diagrams for comparison. Symbols as in Fig. 2.

962 Supplementary Fig. S7: Fractional crystallization models using thorium (Th) as the most  
963 incompatible element (thus considered as a proxy of a quantitative differentiation index). The  
964 numbers along the curves indicate the fraction of residual liquid (f) starting from the basanite RT-  
965 06I-354B, and the bulk partition coefficients for the element (fitted also by visual guess) in the  
966 *ordinata* are reported in each part of the curves. Ticks on the curves are at f=0.9, 0.8, 0.7, 0.6,  
967 0.5, 0.45, 0.4, 0.35, 0.3, 0.25 and 0.2, starting from basanite RT-06I-354B (unreported for  
968 clarity). Symbols as in Fig. 2.

969 Supplementary Fig. S8: a) REE patterns of the Itasy mafic rocks; b) the REE patterns of the mineral  
970 phases of RT-06I-355 and the host rock; c), d) REE patterns of the evolved rocks with c) model  
971 obtained by subtraction of 2% apatite and 1% titanite from sample RT-06I-387, by using the

972 partition coefficients of Fedele et al. (2015) and Olin and Wolff (2012); and d) model obtained  
973 after subtraction of 0.7 and 1% titanite from sample RT-06I-387; e) The REE concentration of  
974 the Itasy rocks normalized to that of the most primitive sample (RT-06I-354B). The effects of  
975 titanite and apatite removal (cf. figure c) are also shown; f) Ratios between the concentrations of  
976 the average composition of titanite, amphibole and clinopyroxene and that of the host rock. The  
977 experimental partition coefficients of titanite in phonolites (Olin and Wolff, 2012),  
978 clinopyroxene in trachyphonolites (Fedele et al., 2009), and amphibole in trachytes (Tiepolo et  
979 al., 2007) are reported for comparison (data in Supplementary Table 11). g) The REE patterns of  
980 the Itasy titanites and those of alkaline intrusive rocks of within-plate affinity (data from Marks  
981 et al., 2008 and Melluso et al., 2010). Prometium is always interpolated. The chondrite is the CI  
982 of Boynton (1984).

1 **The magmatic evolution and genesis of the Quaternary basanite-trachyphonolite suite of Itasy**  
2 **(Madagascar) as inferred by geochemistry, Sr-Nd-Pb isotopes and trace element distribution**  
3 **in coexisting phases**

4  
5 L. Melluso\*<sup>1</sup>, R.D. Tucker\*\*, C. Cucciniello\*, A.P. le Roex\*\*\*, V. Morra\*, A. Zanetti\*\*\*\*, R.L.  
6 Rakotoson\*\*\*\*\*

7  
8 \*DISTAR, Università di Napoli Federico II, Napoli, Italy

9 \*\*BRGM, Orleans, France

10 \*\*\*University of Cape Town, Rondebosch, Republic of South Africa

11 \*\*\*\*IGG-CNR, Pavia, Italy

12 \*\*\*\*\*Analamaity, Antananarivo, Madagascar

13  
14 <sup>1</sup>corresponding author: melluso@unina.it

15  
16 **Abstract**

17  
18 The Itasy is a Pleistocene-Holocene volcanic field in central Madagascar, located to the west of the  
19 Ankaratra volcanic complex. It comprises scoria cones and lava domes (>120), with associated  
20 pyroclastic fall and mafic lava flows, covering an area of ab. 400 km<sup>2</sup>. The last volcanic episodes  
21 probably dated ca. 6000-7100 y BP; warm springs and geysers are active. The juvenile samples  
22 comprise a peculiar, almost bimodal, rock suite ranging from potassic leucite-kaersutite-bearing  
23 basanites, tephrites and phonotephrites, to benmoreites and titanite-haüyne-bearing trachyphonolites  
24 (MgO from 9-10 wt.% to 0.1 wt.%). These rocks show continuous and overlapping variations in the  
25 bulk-rock and phase composition (olivine, clinopyroxene, amphibole, feldspar, leucite, haüyne,  
26 nepheline, oxides, apatite, titanite, glass and other accessories). The basanites have homogeneous  
27 isotopic composition (<sup>87</sup>Sr/<sup>86</sup>Sr=0.70366-0.70378, <sup>143</sup>Nd/<sup>144</sup>Nd=0.51274-0.51277, <sup>206</sup>Pb/<sup>204</sup>Pb=18.7-  
28 18.9, <sup>207</sup>Pb/<sup>204</sup>Pb=15.53-15.56; <sup>208</sup>Pb/<sup>204</sup>Pb=38.89-39.01), and a marked enrichment in the most  
29 incompatible elements (LILE and HFSE reach 100-215 times primitive mantle). These features are  
30 consistent with low degrees of partial melting of a volatile-, LILE- and HFSE-rich, amphibole-  
31 bearing peridotitic mantle induced by uplift during an E-W-directed extensional regime, as is found  
32 in central Madagascar. The marked changes in the geochemical composition, and small variations of  
33 the Sr-Nd-Pb isotopes in the trachyphonolites (<sup>87</sup>Sr/<sup>86</sup>Sr=0.70425-0.70446, <sup>143</sup>Nd/<sup>144</sup>Nd=0.51266-  
34 0.51269, <sup>206</sup>Pb/<sup>204</sup>Pb=18.18-18.39, <sup>207</sup>Pb/<sup>204</sup>Pb=15.49-15.51; <sup>208</sup>Pb/<sup>204</sup>Pb=38.38-39.57) with respect  
35 to basanites and tephrites point to a limited amount of crustal contamination by the relatively low-

36  $^{206}\text{Pb}/^{204}\text{Pb}$ , low- $^{143}\text{Nd}/^{144}\text{Nd}$ , high- $^{87}\text{Sr}/^{86}\text{Sr}$  Precambrian basement rocks (of Middle Archean to Late  
37 Proterozoic age), and highlight the geochemical effects of titanite and anorthoclase removal on the  
38 trace element fractionation trends, a feature also shown in the trace element composition of the  
39 phenocrysts in the trachyphonolites.

40

41 Keywords: basanites, trachyphonolites, mineral trace elements, titanite removal, crustal  
42 contamination, Itasy, Madagascar

43

## 44 **1. Introduction and Geological setting**

45

46 Madagascar is the site of abundant Cenozoic-to-Recent volcanism (Woolley, 2001; Tucker et al.,  
47 2008; Cucciniello et al., 2011, 2016, 2017), which crops out from the northernmost regions of the  
48 island (Cap d'Ambre/Bobaomby, Massif d'Ambre) to the southwest (Ankililoaka, Tulear region;  
49 Fig. 1a). This volcanism is the response to extensional processes and uplift active since the Cenozoic  
50 of the Madagascan lithosphere, with changing orientation of the stress pattern in different sectors of  
51 the island (Nicollet, 1984; Piqué et al., 1999; Cucciniello et al., 2016, 2018). The changing ages and  
52 lithological characteristics of the exposed crustal domains are reflected in the different isotopic  
53 composition of the mantle-derived Cenozoic rocks, suggesting a strong lithospheric control in the  
54 genesis of the latter (Tucker et al., 2014; Melluso et al., 2016; Cucciniello et al., 2018).

55 The central field of alkaline igneous rocks consist of four different volcanic-dominated massifs,  
56 from east to west: Takarindiona, Alaotra, Ankaratra and Itasy (Fig. 1a). The eastern massifs of  
57 Takarindiona and Alaotra are relatively small, structurally controlled Miocene-Pliocene fissural  
58 eruptions and spatter cones (Melluso et al., 2011a), emplaced along or near the western margin of  
59 the oldest rocks in Madagascar (~3.2 Ga gneisses of the Antongil/Masora domain). The Ankaratra  
60 and Itasy are significantly larger in size and more diverse in morphology, as they include lava  
61 domes, massive flow eruptions, spatter cones, and various phreatomagmatic rings and cones (cf.  
62 Brenon and Bussiere, 1959; Woolley, 2001; Cuciniello et al., 2017). The Ankaratra covers an area of  
63 ~3600 km<sup>2</sup> and reaches 2395 mt. elevation. Like the eastern fields, the age of activity in the  
64 Ankaratra spans from Miocene to Recent (Tucker et al., 2008; Rufer et al., 2014; Cucciniello et al.,  
65 2017).

66 The Itasy volcanic field is a much smaller (roughly 400 km<sup>2</sup>) and younger version of the Ankaratra;  
67 it crops out at roughly 1200 mt. a.s.l., in a north-south-trending graben located in the Precambrian  
68 basement (e.g., Andrianaivo and Ramasianoro, 2010, 2011), whose western limits are best indicated  
69 by a long NNE-SSW normal fault (Fig. 1b). The volcanic field is the coalescence of more than 120  
70 scoria cones (e.g., Kassigie, also known for a presumed eruption in 2001, likely a landslide in its SW

71 sector), with lava flows (e.g., along the western master fault), lava domes (e.g., Ngilofotsy, west of  
72 Analavory; Fig. 1b; Supplementary Fig. 1), and pyroclastic rocks in the facies of pyroclastic fall  
73 deposits, small-scale pyroclastic flows and maar craters (cf. Supplementary Fig. 1; Razafiniparany et  
74 al., 1974). According to Bussiere (1957, and references therein), the sequence of volcanic events is  
75 established by cross-cutting relationships. It includes: (a) oldest eruptions of trachyte and  
76 trachyphonolite, (b) deposition of lacustrine sediments bearing fossils of Pleistocene or younger age  
77 (i.e. plant fragments, and bones of hippos, lemurs, and *Aepyornis*), (c) lava flows and pyroclastic  
78 eruptions of basanite and tephrite, and (d) recent strombolian and phreatomagmatic eruptions of  
79 basanitoids. Given the small volume, the scattered nature of magmatism and the structural  
80 complexity of this part of Madagascar, the eruptive history of the Itasy area is only generally known.  
81 What is clear, however, is that the age of the volcanism must be young, certainly less than 1 Ma.  
82 This is consistent with the geologic youthfulness of the volcanic landforms, their relationship to  
83 Pleistocene or younger fossils, the historical record of seismic and geothermal activity (Brown et al.,  
84 2015), and an  $^{40}\text{Ar}/^{39}\text{Ar}$  plateau age of  $97\pm 6$  ka for a trachyphonolite sampled near Analavory (R.D.  
85 Tucker, unpub. data). During and after formation of the lacustrine basins, the region is believed to  
86 have experienced regional E-W extension and differential uplift that fractured the natural dams,  
87 drained the tectonic lakes, and produced its rugged topography. The bounding faults of the Itasy  
88 graben (Fig. 1b) are but one manifestation of this near-surface extensional stress field. These  
89 features, including the volcanism of Itasy and Ankaratra, are a response to still unresolved forces,  
90 recently imaged in a regional seismological survey (Pratt et al., 2017; Fig. 1c) that acted throughout  
91 the Malagasy lithosphere in Miocene and younger time.

92 The Itasy and the Ankaratra are built upon a Precambrian basement complex of Neoproterozoic granitic  
93 gneiss (2.7-2.5 Ga, Betsiboka Suite) and Neoproterozoic quartzite and schist (<1.0 Ga,  
94 Ambatolampy Group), and both are bounded by faults near the apex of the Taná virgation, where the  
95 gneissic fabrics turn sharply from N-S to E-W strike directions (Roig et al., 2012; Tucker et al.,  
96 2014).

97 Petrogenetic studies on the Ankaratra volcanic rocks were recently published (Melluso et al., 2016;  
98 Cucciniello et al., 2017). This paper is focused on the petrogenesis of the Itasy, probably the least  
99 known Late Cenozoic *volcanic* field in Madagascar. Lacroix (1923) and Bussiere (1957) first  
100 identified the featuring petrographic characteristics of the main lithologies, including the presence of  
101 h a y ne in the felsic samples, allowing them to make comparison to rocks of the Massif Central in  
102 France. The range of chemical and isotopic composition of bulk rocks and mineral phases, the  
103 magmatic processes which acted in the upper crust, and the petrogenetic relationships with the other  
104 Cenozoic volcanic districts of Madagascar, particularly those of the central Madagascar, will be  
105 highlighted here.

106

## 107 2. Classification and petrography

108

109 Sixty-nine samples were collected throughout the volcanic field (Fig. 1b). They are representative of  
110 scoria and lava flows cropping out in the graben, and of the several volcanic domes of evolved  
111 composition (trachytic *s.l.*), which occur in the same areas as scoria cones and lavas. The analytical  
112 techniques are fully described in the appendix. According to petrography, T.A.S., R<sub>1</sub>R<sub>2</sub> classification  
113 diagrams (Fig. 2), and CIPW norms (Supplementary Table S1), the fresh samples are basanites,  
114 tephrites, rare phonotephrites and benmoreites (*trachyandesites*), and trachyphonolites, with a  
115 bimodal character (Fig. 2, Table 1 and Supplementary Table S1). The mafic rocks have potassic  
116 affinity, given that their Na<sub>2</sub>O/K<sub>2</sub>O ratios are close to the unity; however, CIPW norms have no  
117 normative leucite. The trachyphonolites do not reach fully peralkaline conditions [P.I., peralkaline  
118 index, molar (Na+K)/Al <0.98], with the exception of the groundmass of evolved samples (see  
119 below). The samples with L.O.I. values > 2 wt.% (mostly scoria) have signs of chemical alteration,  
120 and thus were excluded from the discussion and from the classification diagrams. A synopsis of the  
121 observed mineral assemblages for various lithotypes is reported in the Supplementary Table S2.

122 **Basanites, tephrites and phonotephrites** are variably porphyritic (in both lava and scoria facies),  
123 with zoned, Ti-rich purple-green clinopyroxene, olivine (with small Cr-spinel inclusions), amphibole  
124 and oxide phenocrysts and megacrysts in a fine grained, feldspar-rich mesostasis, to almost aphyric  
125 facies, with purple-green clinopyroxene, oxide and olivine microlites in fine-grained or glassy  
126 groundmass, often without amphibole. The latter phase is found in both scorias and lavas  
127 (Supplementary Fig. S1). Amphibole is almost always poikilitic over clinopyroxene (Supplementary  
128 Fig. S1), and may show reaction relationships with the latter; it is often replaced by secondary  
129 rhönite. There is complete lack of feldspar phenocrysts. Apatite microphenocrysts and microlites are  
130 ubiquitous, particularly in the phonotephrite RT-06I-398, and leucite is often present in the  
131 groundmass, together with haüyne, phonolitic glass and alkali feldspar (Supplementary Fig. S1;  
132 Supplementary Table S2).

133 **Benmoreites**; sample RT06I-387 (the “*ordanchite*” of Lacroix and Bussiere) has rare phenocrysts of  
134 plagioclase, clinopyroxene, amphibole and opaque oxides, set in a trachytic/aphanitic mesostasis  
135 (Supplementary Fig. S1).

136 **Trachyphonolites** range from porphyritic to nearly aphyric, and are characterized by phenocrysts of  
137 anorthoclase and/or sanidine, plus rarer titanite, green clinopyroxene, amphibole and opaque oxides,  
138 set in a trachytic mesostasis that also has apatite, nepheline, haüyne, sodalite, clinopyroxene, other  
139 accessories and interstitial glass (Supplementary Fig. S1; Supplementary Table S2).

140

### 141 3. Mineral compositions

142

143 **Olivine** in basanites and tephrites is usually zoned, with composition ranging from Fo<sub>81</sub> to Fo<sub>55</sub> [Fo  
144 is 100 Mg/(Mg+Fe) in atoms], with MnO up to 0.52 wt.% and CaO up to 0.73 wt.% in the most Fe-  
145 rich compositions (Supplementary Table S3). Some crystals appear too Mg-rich to be in equilibrium  
146 with the host rock composition [e.g., the rounded Fo<sub>81</sub> (xeno)crystal in the RT-06I-398  
147 phonotephrite; Supplementary Table S3].

148 **Oxides. Cr-spinel** is rare and chemically zoned in the Mg-rich olivine cores of basanite RT-06I-  
149 354B [Cr<sub>2</sub>O<sub>3</sub> =8.7-28.7 wt.%; Cr#=47-57; Cr# is 100Cr/(Cr+Al) in atoms]. **Ilmenite** is also very  
150 rare; it was found in the basanite RT-06I-354b as an exsolved phase (Supplementary Fig. S1) and is  
151 rich in geikielite component (MgO=7.8-13.3 wt.%), consistent with its crystallization from a mafic  
152 melt. **Titaniferous magnetite** is the main oxide; it shows increase in MnO and magnetite  
153 component (*s.s.*) and decrease in TiO<sub>2</sub>, Al<sub>2</sub>O<sub>3</sub> and MgO from mafic to evolved rocks (Supplementary  
154 Table S4 and Supplementary Fig. S2).

155 **Clinopyroxene** ranges from diopside [Mg#=82, TiO<sub>2</sub>=1.3 wt.%; Mg# is 100Mg/(Mg+Fe)] to  
156 titanaugite (Mg#=68; TiO<sub>2</sub>=5.75 wt.%, ca. 19 mol.% CaTiAl<sub>2</sub>O<sub>6</sub>) in basanites and tephrites, and  
157 from titanaugite (Mg#=67; TiO<sub>2</sub>=5.1 wt.%) to aegirinaugite (Mg#=22; Na<sub>2</sub>O=5 wt.%, ca. 38 mol.%  
158 aegirine) in the trachyphonolites (Fig. 3). The composition of clinopyroxene is therefore indicative  
159 of the degree of evolution of the host lithotypes, as shown in the Ti-Al and Mg-Ti diagrams  
160 (Supplementary Fig. S3). The evolved rocks have the lowest Mg#, Ti and Al and the highest Mn  
161 (and Na) concentrations; the calculated Al<sup>VI</sup> is absent or strongly subordinated with respect to Al<sup>IV</sup>  
162 (Supplementary Fig. S3; Supplementary Table S5). The compositional ranges are well within those  
163 already observed in the Ankaratra and in the other Cenozoic volcanic areas of Madagascar (e.g.,  
164 Melluso et al., 2011a; Cucciniello et al., 2017 and references therein).

165 **Amphibole** is potassian kaersutite in basanites and tephrites (TiO<sub>2</sub>=4.1-5.1 wt.%; Mg#=58-72), and  
166 potassian kaersutite/Fe-kaersutite to pargasite (TiO<sub>2</sub>=1.2-4.6 wt.%; Mg#=23-56) in benmoreite and  
167 trachyphonolites (Fig. 3). Amphibole has the highest Mg# in the most Mg-rich lava sample (RT-06I-  
168 354B), and has *systematically* lower Mg# than the coexisting clinopyroxene, confirming its later  
169 appearance in the crystallization order (Supplementary Table S6; Supplementary Fig. S1), and thus  
170 bearing no chemical evidence for its crystallization at mantle depths. As with clinopyroxene, there is  
171 no correlation between proxies of the crystallization pressure (e.g., the calculated Al<sup>VI</sup>) and the  
172 composition of amphibole in its host lithology (Supplementary Fig. S4). **Biotite** (Mg#=70-76;  
173 TiO<sub>2</sub>=6.5-6.8 wt.%) is very rare. **Rhönite** [Ca<sub>2</sub>(Mg,Fe<sup>2+</sup>,Fe<sup>3+</sup>,Ti)<sub>6</sub>(Si,Al)<sub>6</sub>O<sub>20</sub>] is often found as a  
174 (subsolidus) rim of amphibole.

175 **Feldspar.** Tiny microlites of labradorite (An<sub>56-59</sub>) to oligoclase (An<sub>27</sub>) are found in the groundmass  
176 of basanites and tephrites, with additional interstitial anorthoclase (Fig. 4a), whereas a continuous  
177 trend from andesine (An<sub>42</sub>) to anorthoclase and sanidine (up to Or<sub>57</sub>) is found in the trachyphonolites  
178 (Fig. 4a). Sr is higher than Ba in plagioclase; they reach 1.9 and 4 wt.%, as oxides, respectively, in  
179 alkali feldspar grains of trachyphonolites (Supplementary Table S7). The overall chemistry is more  
180 restricted than that observed for the feldspars at the Ankaratra volcanic complex (Fig. 4a).  
181 The interstitial **glass** is trachyphonolitic to phonolitic in composition (Supplementary Table S7).  
182 **Feldspathoids.** **Leucite** has been found in the groundmass of the basanites; **häüyne** is found in  
183 phonotephrites, benmoreites and, particularly, in the groundmass of the trachyphonolites. Häüyne  
184 has a significant compositional range, with an almost continuous increase of Na and Cl and decrease  
185 of Ca, K and S towards the late crystallized **sodalite** microlites, also found as rims of zoned crystals  
186 (Fig. 4b, 4c; Supplementary Fig. S1; Supplementary Table S8). The composition of the volatile-  
187 bearing feldspathoids broadly matches the chemical variation of the Mt. Vulture analogues (which  
188 are likely the most complete series of sodalite-group *magmatic* feldspathoids), with the exclusion of  
189 the most potassic varieties (cf. Fig. 4b, and Melluso et al., 2011b and references therein). A Si-rich  
190 **nepheline** (Q<sub>8-24</sub>K<sub>S5-13</sub>Ne<sub>68-78</sub> in wt.%), with low K<sub>2</sub>O (up to 4 wt.%) and variable CaO (up to 2.8  
191 wt.%) completes the groundmass assemblage of holocrystalline samples, including basanites and  
192 tephrites (Supplementary Fig. S5; Supplementary Table S8). The Itasy nepheline differs most  
193 markedly from that of the Madagascan olivine melilitites, the latter having low silica and very high  
194 K (10-12.6 wt.% K<sub>2</sub>O; Melluso et al., 2011a; Supplementary Fig. 5), though it still occurs as a  
195 groundmass phase.

196 **Titanite** occurs as rare phenocrysts in benmoreite and trachyphonolites (Supplementary Fig. 1) and  
197 modally decreases in the most evolved rocks; it usually co-crystallizes with clinopyroxene. The FeO<sub>t</sub>  
198 varies from 1.6 and 7.4 wt.%, the  $\sum$ REE<sub>2</sub>O<sub>3</sub> range from 1.4 to 5.7 wt.%, and ZrO<sub>2</sub> reaches  
199 concentrations as high as 7.8 wt.%, but it is usually far lower (0-4 wt.%). **Apatite** is ubiquitous as a  
200 groundmass phase, and can be found also as microphenocrysts in the phonotephrites. The  
201 concentration of F varies from 1.4 to 4.9 wt.% Cl can be high (up to 1.8 wt.%; average 1.19 wt.%),  
202 and SO<sub>3</sub> is locally significant (up to 1 wt.% in crystals of a tephrite and 2.3 wt.% in crystals of  
203 phonotephrite RT-06I-398; Supplementary Table S9). Rare crystals of **britholite**  
204 [(Ca,LREE,Th)<sub>5</sub>(SiO<sub>4</sub>,PO<sub>4</sub>)<sub>3</sub>(OH,F)] (SiO<sub>2</sub>=21.7 wt.%;  $\sum$ REE<sub>2</sub>O<sub>3</sub>=52 wt.%, F=1.9 wt.%, no Cl and  
205 SO<sub>3</sub>) are found in the groundmass of trachyphonolites. **Zircon** has been found in the interstices of a  
206 trachyphonolite. **Fluorite** is also noted, as well as **hiordahlite** [ideally  
207 (Ca,Na)<sub>3</sub>(Zr,Ti)Si<sub>2</sub>O<sub>7</sub>(F,O,OH)<sub>2</sub>] in the groundmass of the trachyphonolite RT-06I-379  
208 (Supplementary Table 9). Hiordahlite has been noted in the Ampasindava phonolites (northern  
209 Madagascar; Cucciniello et al., 2016) and, together with the Na-rich clinopyroxene, indicates the



210 transition of the groundmass liquids to peralkaline compositions. Other accessories and secondary  
211 zeolites (mostly analcime around nepheline) are present in the interstices.

212

#### 213 **4. Bulk-rock geochemistry**

214

215 The concentration of MgO, Cr and Ni in the most primitive basanites (9-9.2 wt.%, 277-288 ppm and  
216 154-164 ppm, respectively), the Mg# (56-59) and the relatively low forsterite content of the olivine  
217 phenocrysts indicate that the most primitive Itasy rocks can be considered as evolved liquid  
218 compositions which suffered previous removal of Mg-rich olivine, Cr-spinel and Mg-rich  
219 clinopyroxene from more primitive basanites. At the same time, basanites and tephrites have high  
220 concentration of TiO<sub>2</sub> (3-4.7 wt.%) and P<sub>2</sub>O<sub>5</sub> (0.6-1.9 wt.%) (Table 1; Supplementary Table S1). The  
221 variations of CaO, TiO<sub>2</sub>, Al<sub>2</sub>O<sub>3</sub>, Fe<sub>2</sub>O<sub>3t</sub> and alkalis of the mafic rocks, and the mineral composition  
222 in basanites and tephrites depict trends at least qualitatively compatible with removal of pyroxene,  
223 amphibole, olivine and oxides, and minor or negligible removal of feldspar and apatite  
224 (Supplementary Fig. S5). The transition from phonotephrite to benmoreite and trachyphonolites is  
225 characterized by no significant variation of Al<sub>2</sub>O<sub>3</sub>, a change of slope of many minor and trace  
226 element trends, a drop in the concentration of P<sub>2</sub>O<sub>5</sub>, and a decrease of Y with MgO. The  
227 concentration of Ba and Sr also drops in the most evolved trachyphonolites (Supplementary Fig. S6).  
228 The variations of Rb and Ba of the basanites and other features suggest some heterogeneity in the  
229 composition of the mafic parental magmas. The Zr/Hf ratios increase from 45 in the basanites to 60  
230 in the evolved trachyphonolites, whereas the Nb/Ta ratio is 14-15 in the basanites and 21-28 in the  
231 most evolved trachyphonolites (Table 1; Supplementary Fig. S7). The La/Nb, Ba/Nb, Ce/Pb and  
232 Nb/U ratios of the least evolved basanites (La/Nb=0.83-0.88, Ba/Nb=8-12, Ce/Pb=30-41 and  
233 Nb/U=42-52) match the values expected from mantle-derived magmas free of effects related to  
234 crustal contamination in the source or *en route* to the surface, and the low Zr/Nb of the basanites  
235 (4.2-4.8) indicate a highly enriched source region (Melluso et al., 2016 and references therein). The  
236 REE patterns of the mafic rocks (Supplementary Fig. S8a) are smooth and highly fractionated  
237 (La/Yb<sub>n</sub>=24-27), have no peaks or troughs at Eu, and increase in the total REE concentration (279-  
238 800 ppm) from basanites to the phonotephrites. Starting from the phonotephrites RT-06I-398 and  
239 RT-06I-397, the REE patterns of benmoreites and trachyphonolites become increasingly concave,  
240 with a marked decrease of intermediate-REE to values even lower than those of the basanites  
241 (sample RT-06I-357B; Supplementary Fig. S8a, c, d, e), largely constant La<sub>n</sub> and Lu<sub>n</sub>, and small  
242 troughs at Eu (Eu/Eu\*=0.79). These characteristics are highlighted when REE are compared to the  
243 concentration of the least evolved basanite RT-06I-354B (Supplementary Fig. S8e). The mantle  
244 normalized patterns of the Itasy basanites (Fig. 5a, 5b) are almost flat from Rb to Th, have a peak at

245 Nb and Ta, a trough at K and Pb, and a smooth decreasing slope towards the HREE, with no trough  
246 at Zr and Hf, confirming the similarity of the trace element enrichment in the mantle sources  
247 throughout the Cenozoic province (cf. Melluso et al., 2016). The very marked geochemical  
248 differences with the primitive (MgO=9-12 wt.%) *haiiyne-leucite*-basanites of Mt.Vulture (Italy),  
249 which were generated in a subduction-modified upper mantle source (e.g., Beccaluva et al., 2002),  
250 and the similarities with the potassic basanites of the Virunga area (Congo-Rwanda, western branch  
251 of the East African Rift), another within-plate volcanic area as is Itasy (e.g., Rogers et al., 1992), are  
252 highlighted in Fig. 5.

253

## 254 **5. The concentration of trace elements in titanite, clinopyroxene, kaersutite and anorthoclase** 255 **of Itasy trachyphonolites**

256

257 The mantle normalized patterns of the evolved rocks (Fig. 5c) show relative decrease of P, Ti, Ba,  
258 Sr, Ta, Y and middle-REE, and relative increase of Rb, Th, U, Nb, Zr and Hf from the least to the  
259 most evolved trachyphonolites. In order to understand the changes of the chemical composition of  
260 the trachyphonolites, we also investigated the geochemistry of titanite, kaersutite, clinopyroxene and  
261 anorthoclase phenocrysts of the trachyphonolite RT-06I-355 (Table 2; Fig. 6; Supplementary Fig.  
262 S8). The **titanites** have very high concentration of REE ( $\Sigma$ REE=28,000-31,500 ppm), Y (1730-2250  
263 ppm), Ta (650-1130 ppm), Th (134-187 ppm), V (216-328 ppm) and Hf (140-202 ppm), high Zr  
264 (4500-5800 ppm), Nb (6200-15230 ppm), U (15-27 ppm) and La/Yb<sub>n</sub> ratios (19-28), no troughs at  
265 Eu in the chondrite-normalized REE diagrams (Supplementary Fig. S8b), and low Ba (42-49 ppm),  
266 Pb (<1.1 ppm) and Sr (54-382 ppm). The Nb/Ta ratio of the titanites is low (8-13) and the Th/U is  
267 high (8.5-10.3), with respect to typical chondritic ratios (15 and 3.6-4, respectively). The titanites  
268 also have slight chemical zoning, with an increasingly concave REE pattern (Table 2;  
269 Supplementary Fig. S8b). The patterns of the Itasy titanites are generally similar, but displaced to  
270 higher REE concentrations, to those reported by Marks et al. (2008) and Melluso et al. (2010) on  
271 titanites of alkaline intrusive rocks of variable degree of evolution (clinopyroxenites, ijolites,  
272 monzosyenites, to agpaitic nepheline syenites; Supplementary Fig. 8g), which also have widely  
273 variable Nb/Ta ratios (8-66). The Itasy titanite/bulk-rock elemental ratios are of the same magnitude  
274 of the partition coefficients of Olin and Wolff (2012) only for La, Ce and Lu (Supplementary Fig.  
275 8f). **Kaersutite** is a corroded phase, and has a different trace element pattern, with a concentration of  
276  $\Sigma$ REE which is one or two orders of magnitude lower than the concentration of REE of the  
277 coexisting titanites (911 ppm; La<sub>n</sub>= 419 times chondrite; Table 2). It still has relatively high Li (144  
278 ppm), Y (81 ppm), Zr (732 ppm), Nb (200 ppm), moderate Ba (1100 ppm) and very low Rb, Pb, Th

279 and U, when compared to the host-rock bulk composition (Table 1; Fig. 6b). The **clinopyroxene** has  
280 low concentration of REE ( $\Sigma\text{REE}=350\text{-}690$  ppm;  $\text{La}/\text{Yb}_n=8\text{-}15$ ), and very low concentration of other  
281 elements (Table 2), excluding Zn (600-662 ppm). **Anorthoclase** has high and variable Ba (1500-  
282 28220 ppm), Sr (534-4300 ppm), Rb (7-98 ppm), Pb (4.6-9 ppm) and Eu (0.5-1.6 ppm). These  
283 elements peak in the mantle-normalized diagrams (Fig. 6b). Other chemical details can be found in  
284 the Table 2.

285

## 286 **6. Isotopic composition of the Itasy rocks**

287

288 The isotopic composition of Itasy rocks is reported in the Table 3. The basanites (MgO= 7-10 wt.%)  
289 have a restricted range of  $^{87}\text{Sr}/^{86}\text{Sr}$  (0.70366-0.70378),  $^{143}\text{Nd}/^{144}\text{Nd}$  (0.51274-0.51277) and Pb  
290 isotopes ( $^{206}\text{Pb}/^{204}\text{Pb}=18.7\text{-}18.9$ ;  $^{207}\text{Pb}/^{204}\text{Pb}=15.53\text{-}15.56$ ;  $^{208}\text{Pb}/^{204}\text{Pb}=38.89\text{-}39.01$ ), which indicate  
291 a largely uniform mantle source, characterized by low time-integrated Rb/Sr and Sm/Nd ratios. This  
292 source is isotopically similar, but not identical, to that of the mafic alkaline rocks of the Ankaratra  
293 (nephelinites, basanites and alkali basalts; cf. Melluso et al., 2016; Cucciniello et al., 2017; Fig. 7)  
294 and to that of the olivine melilitites to the east and northeast of Ankaratra (Melluso et al., 2011a),  
295 which do not have evidence of crustal contamination. The phonotephrites have  $^{87}\text{Sr}/^{86}\text{Sr}=0.70394\text{-}$   
296  $0.70398$ ,  $^{143}\text{Nd}/^{144}\text{Nd}=0.51273\text{-}0.51274$ ,  $^{206}\text{Pb}/^{204}\text{Pb}=18.74\text{-}18.84$ ,  $^{207}\text{Pb}/^{204}\text{Pb}=15.53\text{-}15.54$  and  
297  $^{208}\text{Pb}/^{204}\text{Pb}=38.97\text{-}39.02$ , all values broadly similar (or indistinguishable in some isotope ratios) to  
298 those of the basanites, whereas the benmoreite and the trachyphonolites have higher  $^{87}\text{Sr}/^{86}\text{Sr}$   
299 ( $0.70430\text{-}0.70446$ ) and lower  $^{143}\text{Nd}/^{144}\text{Nd}$  ( $0.51266\text{-}0.512682$ ),  $^{206}\text{Pb}/^{204}\text{Pb}$  ( $18.18\text{-}18.39$ ),  
300  $^{207}\text{Pb}/^{204}\text{Pb}$  ( $15.49\text{-}15.51$ ) and  $^{208}\text{Pb}/^{204}\text{Pb}$  ( $38.38\text{-}38.57$ ) than the mafic rocks. Therefore, the isotopic  
301 composition of the Itasy rocks changes with MgO. A broadly similar isotopic variation between  
302 associated mafic and evolved rocks has been noted in both Cretaceous and Cenozoic volcanic rocks  
303 of Madagascar (cf. Cucciniello et al., 2010, 2013, 2017) and indicates that, if basanites and  
304 trachyphonolites (which crop out in the same area and have the same stratigraphic age) are part of a  
305 comagmatic suite, the evolved rocks suffered interaction with the old, low-Pb/Pb basement rocks  
306 during magmatic evolution in low-pressure magma reservoirs. This part is developed in the  
307 discussion.

308

## 309 **7. Discussion**

310

### 311 **7.1. The unusual transition between basanites/tephrites and trachyphonolites**

312

313 The Itasy volcanic rocks have some interesting petrogenetic features shared with those of other  
314 volcanic areas worldwide: 1) the leucite-bearing basanites recall analogues found in other areas (e.g.,  
315 the Roman Province, the western branch of the East African Rift); 2) the presence of hauyne  
316 ( $\pm$ sodalite)-bearing alkaline volcanic rocks is found in alkaline rocks of various tectonic settings  
317 (e.g., Massif Central: Vatin-Perignon, 1968; Eifel: Wörner and Schmincke, 1984; Tenerife: Bryan,  
318 2006; Mt. Etinde: Nkombou et al., 1995; Mt. Vulture: Beccaluva et al., 2002; Melluso et al., 1996,  
319 2011b, among other places), and indicate abundant sulfur in the oxidized state (and chlorine as  
320 well); 3) trachyphonolites similar to those of Itasy are common evolved lithotype of other major  
321 volcanic areas (e.g., Suswa, Kenya; Phlegrean Fields-Ischia, Italy). These trachyphonolites are the  
322 product of low-pressure fractional crystallization processes, possibly in an open system. From a  
323 more general view, basanites and tephrites (and maybe also some nephelinite types) are associated  
324 with phonolites *s.s.*, rather than trachyphonolites, as is found in northernmost Madagascar and  
325 elsewhere (e.g., le Roex et al., 1990; Ablay et al., 1998; Melluso et al., 2007; Berger et al., 2014). In  
326 contrast, trachytes (trachyphonolites) are typically associated with (potassic or sodic) alkali and  
327 transitional basalts (e.g., Marion Island, le Roex et al., 2012; St. Helena, Kawabata et al., 2010 and  
328 references therein; Ischia-Phlegrean Fields, Fedele et al., 2008; Melluso et al., 2012, 2014; Libya,  
329 Lustrino et al., 2012; Mauritius: Sheth et al., 2003; Ashwal et al., 2016; Ethiopian Rift, Ronga et al.,  
330 2010; Kenya Rift: White et al., 2012).

331 The magmatic evolution of the Itasy basanites and tephrites (the most Mg-rich of them equilibrated  
332 with olivine at temperatures just lower than 1200°C, according to Roeder and Emslie, 1970), is  
333 without any evidence of removal of plagioclase or leucite, because these phases are found  
334 exclusively as tiny groundmass microlites (e.g., Supplementary Fig. 1). Sodic plagioclase appears as  
335 a phenocryst phase in benmoreites; anorthoclase and Na-sanidine (plus rare mafic phases) are the  
336 dominant phenocrysts and groundmass phases in the trachyphonolites. It is thus evident that removal  
337 of Na-feldspars was essential for the evolution of (sodic) trachyphonolitic liquids towards the  
338 phonolite minimum. This is also indicated by the position of the evolved samples in the Petrogeny  
339 Residua's System, which plot between the trachytic and the phonolitic minima (slightly less than  
340 900°C, at 1 kbar  $P_{H_2O}$ ), hence within the silica undersaturated field (Fig. 8a, b). The tendency of the  
341 bulk-rock trachyphonolitic compositions to plot close to the composition of their (Na-rich) alkali  
342 feldspars (Fig. 4a) can be considered as the *normal* evolution towards the pertinent minimum melt  
343 compositions. The lack of *phenocrysts* of sodic feldspathoids is another indication that the bulk-rock  
344 compositions approach, but do not reach, the phonolitic minimum. For comparison, the phonolites  
345 (*s.s.*) of northernmost Madagascar (Melluso et al., 2007; Cucciniello et al., 2016) have nepheline  
346 phenocrysts, and plot on the alkali feldspar-nepheline cotectics (Fig. 8b).

347 The volatiles that left trace in the Itasy magmatic system are H<sub>2</sub>O (hosted in amphibole, apatite, the  
348 rare biotite, and glass), SO<sub>3</sub> (hosted in the haüyne-sodalite solid solutions and in apatite), Cl (hosted  
349 in sodalite, apatite and glass) and F (hosted in apatite, amphibole, glass, fluorite and other  
350 accessories). No evidence for a significant role of CO<sub>2</sub> in the petrogenesis of the Itasy basanites and  
351 trachyphonolites has been found, as expected by the insignificant maximum quantity of this  
352 compound that can be theoretically dissolved in basanitic melts at low pressure, when compared to  
353 H<sub>2</sub>O (e.g., Shishkina et al., 2014), a quantity even lower in trachytes and phonolites (Webster et al.,  
354 2014). The Itasy basanites and tephrites have significant amounts of kaersutite, indicating that the  
355 magmatic crystallization took place in a H<sub>2</sub>O-rich environment at relatively early differentiation  
356 stages. Kaersutite is unknown in basanites and alkali basalts (and nephelinites too) of the nearby  
357 Ankaratra complex, that have mostly anhydrous mineral assemblages, and occurs as a phenocryst  
358 phase only in a few benmoreites (Cucciniello et al., 2017). Kaersutite *s.s.* crystallizes in mafic,  
359 hydrous, feldspar-bearing (or -normative), high-Ti alkaline magmas, and has a range of  
360 crystallization depths from the uppermost mantle (e.g., Irving and Green, 2008; Pilet et al., 2010) to  
361 shallow conduits or shallow intrusions (e.g., Melluso et al., 2005, 2007); hence this phase is more a  
362 chemical than a barometric indicator (cf. Supplementary Fig. S4). Sulfur was at least partially  
363 partitioned in the S-bearing apatite before reaching saturation as haüyne, in the groundmass of the  
364 phonotephrites and in the evolved rocks; sulfur decreases in the latest crystallized groundmass, as  
365 testified by the presence of sodalite (cf. also Supplementary Fig. S1).

366

## 367 **7.2. Petrogenetic modelling: the essential role of titanite and anorthoclase removal in the** 368 **geochemical evolution of the trachyphonolites, and open-system processes**

369

370 The petrogenetic study of the Itasy volcanic rocks is focused on the genetic relationships between  
371 this bimodal suite of mafic and evolved rocks, with their changing isotopic ratios and the marked  
372 decrease in the degree of silica undersaturation of the evolved rocks. Despite the compositional gap  
373 between mafic and evolved rocks shown in the diagrams, involving a marked change of density,  
374 temperatures and time of storage in shallow reservoirs, we observe the following: 1) the mineral  
375 phases have a continuous compositional change from mafic, through intermediate, to felsic rocks, as  
376 a typical consequence of fractional crystallization processes (see above); 2) the bulk-rock  
377 compositions have the expected changes considering the variety and amount of the phenocryst  
378 phases, thus excluding effects of magma mixing; 3) many basanites and tephrites are rich in  
379 phenocrysts/megacrysts of kaersutite (a phase significantly lower in silica than clinopyroxene but  
380 with similar density), whose significant removal could help increase the concentration of SiO<sub>2</sub> of  
381 residual melts.

382 The variation diagrams (Supplementary Fig. S6) and the petrography of the mafic rocks indicate that  
383 clinopyroxene, amphibole, olivine and magnetite (for Fe and Ti) are probably the best candidates to  
384 form the subtracted mineral assemblages in the transition from basanites to phonotephrites. Among  
385 several models using the major element compositions of bulk-rocks and their phenocryst phases, the  
386 transition from basanites (e.g. RT-06I-354B) to tephrites (RT-06I-395) and benmoreite (RT-06I-  
387 387) can be accounted for by 37% and 56% removal of ultramafic assemblages, respectively. The  
388 transition from the benmoreite RT-06I-387 to the trachyphonolite RT-06I-375B needs 61% removal  
389 of monzonitic assemblages. The total transition from the basanite RT-06I-354B to a trachyphonolite  
390 (e.g., sample RT-06I-375B) was modelled with 71.4% removal of ultramafic assemblages (cf.  
391 Supplementary Table S10). The results we choose have low  $\Sigma R^2$  (values of  $\Sigma R^2 \ll$  or  $<0.5$  can be  
392 considered for modelling, but they are not proof of geological reliability), but the high  
393 kaersutite/clinopyroxene ratio involved in some transitions between mafic rocks is always to be  
394 matched with the actual subtracted assemblages, which are, of course, largely unknown.  
395 Thermodynamic models of magmatic evolution (e.g., MELTS) are not able to account for the  
396 presence of kaersutite, sodalite-group minerals, titanite and a number of other petrographic features  
397 of alkaline rock suites, including the crystallization sequence, the phase composition and their actual  
398 amount. LeMasurier et al. (2011) hypothesized a significant role for kaersutite removal in order to  
399 account for the petrogenesis of evolved silica oversaturated rocks starting from basanite parental  
400 magmas; in our case, we show that this effect is limited to the transition from tephrites to  
401 benmoreites, because the final transition to the most evolved Itasy trachyphonolites cannot involve  
402 significant removal of amphibole (kaersutite is already a minor phenocryst phase in the least evolved  
403 trachyphonolites).

404 A number of observations and models indicate that the Itasy benmoreites and trachyphonolites could  
405 be comagmatic with the basanites, through (also) the effects of kaersutite removal. The trace element  
406 variations using thorium as the most incompatible element were modelled using the Rayleigh  
407 equation (Supplementary Fig. S7). The following features are of interest: a) assuming that Th  
408 behaves in a closed system (an approximation that cannot be guaranteed during AFC processes, see  
409 below) and that is the most incompatible element, the total amount of crystallization needed to reach  
410 the phonotephritic liquids is roughly 60%, and is 75-80% to reach evolved trachyphonolites, starting  
411 from the least evolved basanites; b) the transition from basic to intermediate rocks is characterized  
412 by several major changes in the slope of the trace elements vs. the Th concentration and, therefore,  
413 in their bulk partition coefficients (D); this change in slope takes place at different evolution points  
414 (Supplementary Fig. S6 and S7); c) tantalum (Ta) behaves as a highly compatible element in the  
415 trachyphonolites, thus strongly decouples from the geochemically similar element Nb; roughly  
416 similar fractionation is noted between Hf and Zr, whereas U and Th behave sympathetically (the

417 Th/U ratio ranges between 3.6 and 4.4, both values in basanites); d) lead (Pb) increases more than  
418 that expected from a perfectly incompatible behaviour (modelling requires  $D_{Pb} < 0$ ; Supplementary  
419 Fig. S7), strongly suggesting the introduction of external Pb during crustal contamination processes;  
420 e) the chemical composition of the trachyphonolites is scattered, indicating that the magmatic  
421 evolution took place in independent magma feeder systems, hence with variable distribution,  
422 crystallization and removal of phases.

423 The petrography, the increasingly concave REE patterns of the trachyphonolites, the trace element  
424 concentration of the titanite phenocrysts, and numerical modelling using both published partition  
425 coefficients and the mineral composition determined by LA-ICP-MS (Supplementary Fig. S8)  
426 indicate a major effect of titanite on the fractionation of middle-REE, Y, Ta and other elements,  
427 even though full quantification of the extent of removal of an accessory phase cannot be obtained  
428 through major oxide mass balance calculations or trace element behaviour. The trend to the most  
429 evolved trachyphonolites can be obtained through 0.7-1% removal of titanite (cf. Supplementary  
430 Fig. S8c, d). Removal of small amounts of titanite is commonplace in evolved alkaline rocks (cf. the  
431 extremely similar patterns reported in Ulrych et al., 2006; Lustrino et al., 2012; Melluso et al., 2014,  
432 2017 and references therein), and the effects observed in this work are broadly those expected  
433 utilizing the titanite/melt partitioning data of Olin and Wolff (2012) (Supplementary Fig. S8c,  
434 Supplementary Table S11). Apatite can have a roughly similar behaviour regarding middle-REE  
435 fractionation, although to a lower extent (the partition coefficients of apatite can be lower of a factor  
436 of 3-4; cf. Supplementary Table S11 and Supplementary Fig. S8c), but has minor or negligible effect  
437 on the concentration of Ta, Hf, Zr and Nb. Removal of other phases such as clinopyroxene or  
438 amphibole, in the amount observed in the least evolved trachyphonolites (Supplementary Fig. S1),  
439 cannot have appreciable role in preferentially removing middle- from heavy-REE in evolved melts  
440 (cf. partitioning data of Tiepolo et al. 2007; Fedele et al., 2009; Melluso et al., 2017; this work), and  
441 so is for other accessory phases *absent* at Itasy (e.g., melanite garnet, or significant zircon).

442 The evolved samples of Itasy show compelling evidence for open-system evolution, as seen from the  
443 isotopic variations and elemental evidence above. We performed an AFC model (DePaolo, 1981)  
444 utilizing all the isotopic systematics, pertinent trace elements and compositions of the Madagascan  
445 basement in the area. The results highlight the effects of prolonged fractional crystallization (roughly  
446 70%) and limited, though widely variable, crustal contamination (the values of the  
447 assimilation/cumulus rate,  $-r$ , are  $< 0.3$ ) of known felsic rocks of the Precambrian basement (details  
448 in Fig. 9). The model takes into account the marked changes in the bulk partition coefficients of  
449 many elements at varying levels of magmatic evolution (Supplementary Fig. S8). The  
450 phonotephrites (samples RT-06I-397 and RT-06I-398) still need limited amount of crustal  
451 contamination in order to account for their slightly differing isotopic ratios compared to those of

452 basanites and tephrites. The abrupt change of partition coefficients in the evolved rocks and a limited  
453 change of the parameter “r” (Fig. 9) led to the bimodal isotopic compositions. It is also evident that  
454 the effects of crustal contamination did not act uniformly throughout the Itasy suite, and that other  
455 processes could have been operative in the Itasy feeder system, including silica increase that may be  
456 expected from interaction with the felsic lithologies of the Madagascan basement, and the presence  
457 of alkali basaltic parental magmas (more SiO<sub>2</sub>-rich than a basanite), of which there is no evidence, *to*  
458 *date*.

459

### 460 **7.3. The mantle sources of Itasy in the context of the Cenozoic magmatism of Madagascar**

461

462 The Itasy basanites have Sr-Nd-Pb-isotopic ratios indicating a source geochemically and isotopically  
463 different from that of convecting asthenosphere, which experienced a distinct evolution with time.  
464 This is fully consistent with what is known in the Cenozoic mafic rocks of the central Madagascar,  
465 which were emplaced above Archean crust (Melluso et al., 2011a, 2016; Cucciniello et al., 2017).  
466 Primitive basanitic magmas with the petrographic and geochemical characteristics similar to those of  
467 the Itasy imply low degrees of partial melting (ca. 3%) of a hydrous, amphibole-bearing lherzolite,  
468 and hence at depths and temperatures where *pargasitic* amphibole can be a stable phase of a  
469 peridotite (say, <90-100 km; e.g., Kovacs et al., 2017), as previously suggested by earlier work in  
470 central and northern Madagascar (Melluso and Morra, 2000; Melluso et al., 2016 and references  
471 therein; Rocco et al., 2017), and elsewhere (cf. le Roex et al., 2001; Green and Falloon, 2015). A  
472 relatively shallow depth of the (lithospheric) source region of the Itasy basanites is also indicated by  
473 a rootless low-velocity zone (supposedly melt-, volatile- or light phases-bearing) just beneath the  
474 Moho, extending at depths of 50-80 km (Fig. 1c; Pratt et al., 2017). The E-W-trending extensional  
475 processes active in central Madagascar seem to have favoured moderate uplift of this geochemically  
476 enriched, volatile-bearing lithospheric mantle, bringing it above the solidus and forming the Itasy  
477 low-volume basanitic lava field. The source of the Itasy basanites must be characterized by a  
478 significant enrichment of highly incompatible elements (cf. the low Ba/Nb and Zr/Nb ratios, and the  
479 shape of the mantle normalized patterns), with particular reference to the HFSE, in agreement with a  
480 fully within-plate origin, without subduction-related components or low-pressure crustal  
481 contamination. The source of the Itasy basanites cannot be depleted asthenospheric -or primitive-  
482 mantle, assuming any reasonable degrees of partial melting and considering the isotopic composition  
483 (cf. also Melluso et al., 2016 and Cucciniello et al., 2018, for a regional review). Finally, in terms of  
484 trace element and isotopic composition, the sources of the Cenozoic volcanic rocks of Madagascar  
485 share very little with those giving rise the Cretaceous *tholeiitic* magmatism (cf. Melluso et al., 2001,  
486 2005, 2011a, 2016; Cucciniello et al., 2010, 2013, 2017).



487

## 488 8. Conclusions

489

490 The Itasy volcanic field, the most recently active in Madagascar, is characterized by an association  
491 of high-TiO<sub>2</sub>, potassic basanites/tephrites/phonotephrites and benmoreites/ trachyphonolites, in the  
492 facies of pyroclastic rocks and lava flows, with the highly featuring presence of Ti-rich amphibole in  
493 mafic/intermediate rocks and h aüyne-sodalite *s.s.* in evolved rocks. Recognizing the possibility of  
494 independent basanitic parental magmas and different liquid lines of descent, leucite basanite  
495 magmas evolved through removal of olivine, clinopyroxene, kaersutite and magnetite to reach the  
496 volumetrically minor benmoreitic and trachyphonolitic compositions. These latter rocks are likely  
497 the result of low-pressure fractional crystallization of anorthoclase/sanidine-bearing (*syenitic*)  
498 assemblages and minor crustal contamination of the high-<sup>87</sup>Sr/<sup>86</sup>Sr, low- <sup>143</sup>Nd/<sup>144</sup>Nd, -<sup>206</sup>Pb/<sup>204</sup>Pb, -  
499 <sup>207</sup>Pb/<sup>204</sup>Pb and -<sup>208</sup>Pb/<sup>204</sup>Pb, old granitic-migmatitic basement of the area, with the evidence of  
500 minor titanite removal (perhaps <2%), as part of syenitic/monzosyenitic cumulates. Titanite is  
501 demonstrated to be a very efficient scavenger of intermediate-REE, Y, Ta and V, and slightly less so  
502 for Nb, Zr and Hf in the Itasy trachytic (*s.l.*) magmas (cf. Table 2); its removal, even in small  
503 amounts, significantly increases ratios between otherwise geochemically similar elements (e.g.,  
504 Zr/Hf, Nb/Ta) when compared to those of the associated mafic rocks. Another minor but interesting  
505 result of this study is the evidence that the concentration of rare earths or other strategic elements in  
506 the most highly evolved alkaline magmas cannot be considered a rule working worldwide, due to the  
507 presence of accessory phases on the liquidus, which effectively scavenge these elements in less  
508 evolved melts. *Also* due to the removal of titanite and other accessory phases, the geochemical  
509 characteristics of trachytic (*s.l.*) rocks *worldwide* cannot give unambiguous indication about those of  
510 associated (or parental) mafic magmas, or even to indicate mantle signatures of any kind, including  
511 ratios of elements thought to be highly incompatible. The geochemical and isotopic characteristics of  
512 the leucite-basanites broadly match those of the nearby Ankaratra volcano and of the rest of the  
513 Cenozoic volcanic districts of Madagascar (Melluso et al., 2016), and testify to a mantle source  
514 having marked enrichment in LILE, HFSE and volatile elements (H<sub>2</sub>O, Cl, F, SO<sub>3</sub>), as noted in other  
515 within-plate strongly alkaline rocks worldwide and, intriguingly, also in some subduction-related  
516 alkaline rocks. The broadly E-W extensional processes common in central Madagascar triggered  
517 low-degree partial melting of enriched, volatile-rich, lithospheric mantle, and generated the low-  
518 volume Itasy basanites, without the need to invoke the action of mantle plumes. Enriched  
519 geochemical components of within-plate origin are commonplace in the Madagascan lithospheric  
520 mantle (Melluso et al., 2016; Rocco et al., 2017). Finally, the unusual petrogenetic association  
521 between basanites and trachyphonolites, and the likely significant removal of kaersutitic amphibole

522 in some basanites and tephrites, are of importance in identifying the volatile species of within-plate  
523 alkaline rocks in the mantle sources and during the magmatic evolution.

524 A high-resolution geochronological framework is now needed to establish an accurate age span for  
525 the Itasy, if a significant temporal lapse exists between mafic and evolved rocks, and the  
526 relationships between timing of volcanism and extensional tectonics of central Madagascar.

527

## 528 **Acknowledgements**

529

530 This paper is dedicated to the memory of Fabio Carmine Mazzeo.

531 Alessio Langella, Sergio Bravi, Petrus Le Roux, Roberto de' Gennaro, Vohangy Ratrimo,  
532 Dieudonné Razafimahatratra, Vincenza Guarino, Luigi Franciosi and Lorenzo Fedele are gratefully  
533 thanked for helping us in sampling, data acquisition and comments on an early draft of the  
534 manuscript. The constructive reviews of Jean-Paul Liégeois and Michael Marks and advice of the  
535 Editor-in-chief Xian-Hua Li were very helpful to improve contents and presentation of an early  
536 version of the manuscript. This paper was supported by grants provided by PRIN2015 (grant  
537 20158A9CBM) and LR5 (Regione Campania) to L. Melluso, and Fondi Ricerca di Ateneo  
538 (DR\_3450\_2016) to C. Cucciniello.

539

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777

#### 778 **Appendix: analytical techniques**

779

780 Powders of the 69 samples of this study were obtained after carefully grinding washed chips in an  
781 ultrapure agate mill. For each sample, four grams of micronized powder were used to prepare pressed  
782 powder pellets. The powder (mixed with 1 ml of Polyvinyl alcohol solution) was pressed to twenty  
783 tons for twenty seconds. The bulk-rock X-Ray Fluorescence compositional data (Table S1) were  
784 obtained on pressed powder pellets with an Axios Panalytical spectrometer equipped with six analyzer  
785 crystals, three primary collimators and two detectors (flow counter and scintillator), operating at  
786 different kV and mA for each analyte. Analytical uncertainties are in the order of 1-2% for major  
787 elements and 5-10% for trace elements. The weight loss on ignition has been obtained with  
788 gravimetric techniques, firing at 1000°C small aliquots of powders previously dried at 110°C  
789 overnight.

790 Data on a subset of samples were obtained through ICP-MS methods at Actlabs (Canada) (Table 1).  
791 Samples were mixed with a flux of lithium metaborate and lithium tetraborate and fused in an  
792 induction furnace. The molten material is immediately poured into a solution of 5% nitric acid  
793 containing an internal standard, and mixed continuously until completely dissolved (~30 minutes). The  
794 samples were run for major oxides and selected trace elements (Ba, Be, Sc, Sr, V, Y and Zr) on a  
795 combination simultaneous/sequential Thermo Jarrell-Ash ENVIRO II ICP or a Varian Vista 735 ICP.  
796 Calibration is performed using seven prepared USGS and CANMET certified reference materials. One

797 of the seven standards is used during the analysis for every group of ten samples. Sample fused are  
798 diluted and analyzed by Perkin Elmer Sciex ELAN 6000, 6100 or 9000 ICP/MS for other trace  
799 elements (Cr, Co, Ni, Cu, Zn, Ga, Ge, As, Rb, Nb, Mo, Ag, In, Sn, Sb, Cs, La, Ce, Pr, Nd, Sm, Eu,  
800 Gd, Tb, Dy, Ho, Er, Tm, Yb, Lu, Hf, Ta, W, Tl, Pb, Bi, Th and U). Three blanks and five controls  
801 (three before the sample group and two after) are analyzed per group of samples. Duplicates are fused  
802 and analyzed every 15 samples. Analyses of international standards are reported in Cucciniello et al.  
803 (2017).

804 Microprobe analyses were carried out on polished thin sections using Energy Dispersive Spectrometry  
805 (EDS) at University of Napoli Federico II, utilizing a JEOL JSM-5310 microscope operating at 15 kV  
806 primary beam voltage, 50-100 mA filament current, 50 s net acquisition time and a Oxford  
807 Instruments Microanalysis Unit, equipped with an INCA X-act detector. For further details see  
808 Melluso et al. (2014, 2017).

809 The Laser Ablation - Inductively Coupled Plasma - Mass Spectrometry analysis of the minerals in the  
810 trachyphonolite RT-06I-355 was performed at IGG-CNR (Pavia, Italy), utilizing techniques described  
811 in Rocco et al. (2017). Concentrations of Rare Earth Elements (REE) and other selected trace elements  
812 (Li, B, Ba, Rb, Th, U, Nb, Ta, Sr, Zr, Hf, Ti, Y, Sc, V, Co and Ni) were determined in situ on polished  
813 sections (100  $\mu\text{m}$ ) using a PerkinElmerSCIEX ELAN DRC-e quadrupole mass spectrometer coupled  
814 with an UP213 deep-UV YAG Laser Ablation System (New Wave Research, Inc.). The laser was  
815 operated at a repetition rate of 10 Hz, with 213 nm wavelength and a fluence of  $\sim 9.5 \text{ J/cm}^2$ . Helium  
816 was used as carrier gas and was mixed with Ar downstream of the ablation cell. Spot diameter was 55  
817  $\mu\text{m}$ . Data reduction was performed offline using the GLITTER software. For this study, the NIST  
818 SRM 610 synthetic glass standards was used as external standard, CaO was used as internal standard  
819 for clinopyroxene and amphibole, while  $\text{SiO}_2$  was used for titanite and feldspar. Precision and  
820 accuracy of the REE concentration values were assessed through repeated analysis of the BCR2-g  
821 standard to be better than  $\pm 7\%$  and  $\pm 10\%$ , respectively, at the ppm concentration level (for further  
822 details see Miller et al., 2012 and references therein).

823 The Sr-Nd-Pb-isotope data were obtained at the University of Cape Town, with techniques described  
824 in le Roex et al. (2012). Sr, Nd and Pb were separated using conventional ion exchange techniques and  
825 all radiogenic isotope analyses were performed on a NuPlasma multicollector inductively coupled  
826 plasma-mass spectrometer (ICP-MS). To correct for mass fractionation effects, measured  $^{87}\text{Sr}/^{86}\text{Sr}$  and  
827  $^{143}\text{Nd}/^{144}\text{Nd}$  values were normalized to  $^{86}\text{Sr}/^{88}\text{Sr}=0.1194$  and  $^{146}\text{Nd}/^{144}\text{Nd}=0.7219$ , respectively. Pb  
828 isotopes were corrected for fractionation by normalizing ratios measured in international standards.  
829 Average standard values obtained during the course of this study are reported in the Table 3.

830

831 **Table captions**

832

833 Table 1: ICP-MS major and trace element composition of Itasy rocks. A few element ratios are also  
834 reported.

835 Table 2: LAM-ICP-MS and microprobe composition of titanite, amphibole, clinopyroxene and  
836 feldspar of the trachyphonolite RT-06I-355. The location of the laser pits is shown in the  
837 Supplementary Fig. S1c.

838 Table 3: Isotopic composition of the Itasy rocks. The composition of the standards which were run  
839 with the unknowns is also reported.

840

### 841 **Supplementary Tables**

842

843 Supplementary Table S1: XRF major trace element data and CIPW norms of the Itasy samples. The  
844 composition of international standards is also reported.

845 Supplementary Table S2: synopsis of the mineral assemblages of the main lithotypes of the Itasy  
846 complex.

847 Supplementary Table S3: composition of olivine of the Itasy rocks.

848 Supplementary Table S4: composition of oxides of the Itasy rocks.

849 Supplementary Table S5: composition of pyroxene of the Itasy rocks.

850 Supplementary Table S6: composition of amphibole, biotite and rhönite of the Itasy rocks.

851 Supplementary Table S7: composition of feldspar and glass of the Itasy rocks.

852 Supplementary Table S8: composition of feldspathoids of the Itasy rocks.

853 Supplementary Table S9: composition of titanite, apatite, other accessories of the Itasy rocks.

854 Supplementary Table S10: Recapitulation of mass balance calculations between rocks of different  
855 degree of magmatic evolution, and detailed results. The composition of the phases is reported in  
856 the supplementary tables.

857 Supplementary Table S11: average REE mineral/bulk rock ratios for titanite, amphibole and  
858 clinopyroxene. The partition coefficients of titanite, clinopyroxene, apatite and amphibole are  
859 taken from Olin and Wolff (2012), Fedele et al. (2009, 2015) and Tiepolo et al. (2007). Note that  
860 amphibole of RT-06I-355 could not be considered in equilibrium with the host rock (cf.  
861 Supplementary Fig. S1).

862

### 863 **Figure captions**

864

865 Figure 1: a) Simplified geological map of Madagascar, with the location of the main volcanic areas  
866 of Cretaceous and Cenozoic; b) Geological sketch map with sample localities (Laborde

867 coordinates) and the main tectonic features of the Itasy area. c) Schematic section of the Itasy  
868 region (redrawn after Pratt et al., 2017) with location of a possible mantle source region of the  
869 Itasy mafic volcanic rocks.

870 Figure 2: Classification of the Itasy rocks according to the TAS and  $R_1R_2$  diagrams. Symbols: Itasy  
871 mafic rocks: red triangles; Itasy evolved rocks: blue circles. The data from other Cenozoic  
872 districts of Madagascar are taken from: Melluso et al. (2007, 2011b); Cucciniello et al. (2011,  
873 2016, 2017); Melluso and Morra (2000).

874 Figure 3: Classification of pyroxene and amphibole of the Itasy rocks. Symbols as in Fig. 2. The  
875 pyroxene and amphibole composition of the Ankaratra alkaline rocks (black squares) are taken  
876 from Cucciniello et al. (2017).

877 Figure 4: a) The composition of feldspars and bulk-rocks of the Itasy in the An-Ab-Or diagram  
878 (mol.%). b) Composition of the Itasy volatile-rich feldspathoids in the CaO-Na<sub>2</sub>O-K<sub>2</sub>O diagram  
879 (mol.%; after Lessing and Grout, 1971). c) The volatile concentration of the Itasy feldspathoids  
880 (in wt.%). The feldspar composition of the Ankaratra alkaline rocks (black squares) are taken  
881 from Cucciniello et al. (2017). The data on Mt. Vulture feldspathoids are taken from Melluso et  
882 al. (1996, 2011a), references therein and unpublished data. Symbols as in Fig. 2.

883 Figure 5: a, b) Mantle normalized diagrams for the Itasy basanites. The data on the other  
884 Madagascan basanites (grey patterns in the background) are taken from the database of Melluso  
885 et al. (2016). The analyses of the Mt. Vulture basanites (MgO=9-12 wt.%) are taken from  
886 Beccaluva et al. (2002), L.M. and V.M., unpublished data. The data on the Karisimbi leucite  
887 basanites (green patterns in the background) are taken from Rogers et al. (1992). At a similar  
888 MgO, the mantle-normalized patterns of the Mt. Vulture and Itasy basanites (as well other  
889 geochemical characteristics) imply contrasting enrichment processes in their respective sources,  
890 and a different mantle baseline preceding metasomatic events; c) Mantle normalized diagrams  
891 for the Itasy evolved rocks. The normalization values are taken from Lyubetskaya and Korenaga  
892 (2007).

893 Figure 6: a, b) Mantle normalized diagrams of the coexisting phases of the trachyphonolite RT-06I-  
894 355, together with those of the Itasy evolved rocks. Normalization values are those of  
895 Lyubetskaya and Korenaga (2007). Potassium and Phosphorus are not reported.

896 Figure 7: Isotopic diagrams for the Itasy and the other Cenozoic volcanic rocks emplaced within the  
897 cratonic domains of the Madagascan basement (cf. Fig. 1a). Symbols as in Fig. 2. The Ankaratra  
898 data are taken from Cucciniello et al. (2017), and the Alaotra and Takarindiona olivine melilitites  
899 (black triangles) are taken from Melluso et al. (2011a).

900 Figure 8: a) The tetrahedral phase diagram diopside-nepheline-kalsilite-silica at one atmosphere  
901 (after Cucciniello, 2016) with the Itasy bulk-rock compositions (symbols as in Fig. 2); b) The

902 Petrogeny Residua's System at 1 kbar  $P_{H_2O}$  with the Itasy trachyphonolites (blue circles). The  
903 boundary lines are taken from Gupta (2015). The evolved rocks of the Ankaratra (black squares)  
904 are taken from Cucciniello et al. (2017); the Bobaomby phonolites (black triangles) are taken  
905 from Melluso et al. (2007); the trachytes of Mauritius (grey rhombs) are taken from Ashwal et al.  
906 (2016). Other evolved rocks can be found in Cucciniello et al. (2017). The compositions of glass  
907 of the evolved rocks is reported as light blue circles.

908 Figure 9: AFC modelling after DePaolo (1981). The composition of the contaminants (granitic rocks  
909 and felsic gneisses of the basement around Itasy: data from Kröner et al., 2000; Melluso et al.,  
910 2001; Cucciniello et al., 2017) is the following:  $^{87}Sr/^{86}Sr=0.7113$ ,  $^{143}Nd/^{144}Nd=0.5113$ ,  
911  $^{206}Pb/^{204}Pb=16.62$ ,  $Pb=10$  ppm;  $Sr=499$  ppm;  $Nd=16$  ppm;  $Th=12$  ppm. The bulk partition  
912 coefficients are reported along the curves (cf. Supplementary Fig. S7), together with the values of  
913 the residual liquid fraction (f) and the assimilation/accumulus rate (r). The AFC models in the  
914 diagram  $^{143}Nd/^{144}Nd$  vs.  $Nd$  were performed taking into account the geochemical behaviour of  $Nd$   
915 shown in Supplementary Fig. S7, and two starting compositions (the basanite RT-06I-354B and  
916 the benmoreite RT-06I-387). Symbols as in Fig. 2.

917

## 918 **Supplementary Figures**

919

920 Supplementary Fig. S1a: Petrographic, BSE (backscattered electron images), and field images of the  
921 Itasy rocks: a) basanite RT-06I-354B: phenocrysts of olivine, with Cr-spinels, set in a scoriaceous  
922 mesostasis; b) basanite RT-06I-354B: large phenocrysts of kaersutite enclosing diopsidic  
923 clinopyroxene, parallel nicols; c) phonotephrite RT-06I-398: rare phenocrysts of clinopyroxene  
924 and apatite set in a very fine-grained microlitic mesostasis, parallel nicols; d) basanite RT-06I-  
925 354B: large phenocryst of diopside-to-titanaugite clinopyroxene, parallel nicols; e)  
926 trachyphonolite RT-06I-355; phenocrysts of a zoned anorthoclase and titanite in a fine-grained  
927 feldspar-rich mesostasis, parallel nicols; f) trachyphonolite RT-06I-375: phenocrysts of  
928 (pleochroic) titanite and alkali feldspar in a fine-grained groundmass, parallel nicols. The  
929 horizontal scale of the thin section photomicrographs is 2.5 mm.

930 Supplementary Fig. S1b: g) groundmass of basanite RT-06I-354B: microlites of feldspar,  
931 feldspathoids, unmixed Fe-Ti oxides and clinopyroxene, BSE; h) trachyphonolite RT-06I-361:  
932 phenocrysts of amphibole, and microphenocrysts of titanite and apatite in a fine-grained  
933 mesostasis; i) trachyphonolite RT-06I-363: microphenocrysts of Fe-Ti oxides, titanite, and  
934 clinopyroxene in a fine-grained mesostasis rich in alkali feldspar and light grey glass, BSE; j)  
935 phonotephrite RT-06I-398: a microlite of h a yne (darker grey) in the center of the groundmass  
936 made up of feldspar, clinopyroxene, magnetite and glass, BSE; k) l) m) n) o) feldspathoids in the

937 groundmass of trachyphonolites RT-06I-361 and RT-06I-379; p), q), r) pyroclastic rocks, lava  
938 domes, and spatter cones in the northeastern part of Itasy.

939 Supplementary Fig. S1c: Petrographic characteristics of the trachyphonolite RT-06I-355. a) zoned  
940 green cpx phenocryst b), c) phenocrysts of alkali feldspar and titanite; d) microlites of alkali  
941 feldspar and altered amphibole. The laser pits after LAM-ICP-MS work (Table 2) can be easily  
942 seen. The horizontal scale of the thin section photomicrographs is 2.5 mm.

943 Supplementary Fig. S2: Composition of the titaniferous magnetite of the Itasy rocks in the MgO-  
944 MnO and TiO<sub>2</sub>-Al<sub>2</sub>O<sub>3</sub> diagrams (in wt.%). Symbols as in Fig. 2.

945 Supplementary Fig. S3: Clinopyroxene composition. Symbols: mafic rocks: red triangles;  
946 benmoreites and trachyphonolites: blue circles. The composition of olivine is plotted in the Ca-  
947 Mg-Fe diagram.

948 Supplementary Fig. S4: Amphibole composition (cations in apfu) indicating the *absence of evidence*  
949 for a high-pressure crystallization environment for this phase, as the calculated Al<sup>VI</sup> is roughly  
950 1:10 the calculated Al<sup>IV</sup> in both mafic and evolved rocks. Symbols as in Fig. 2

951 Supplementary Fig. S5: composition of nepheline in the nepheline-kalsilite-silica diagram. The  
952 nephelines of the olivine melilitites and of the Ankaratra rocks are reported for comparison (data  
953 from Melluso et al., 2011a and Cucciniello et al., 2017).

954 Supplementary Fig. S6: XRF major oxides and trace elements vs. MgO wt.% (plotted on a  
955 logarithmic scale). The composition of olivine, amphibole and clinopyroxene of the basanites  
956 and tephrites is reported in a few diagrams for comparison. Symbols as in Fig. 2.

957 Supplementary Fig. S7: Fractional crystallization models using thorium (Th) as the most  
958 incompatible element (thus as a proxy of a quantitative differentiation index). The numbers along  
959 the curves indicate the fraction of residual liquid (f) starting from the basanite RT-06I-354B, and  
960 the bulk partition coefficients for the element (fitted also by visual guess) in the *ordinata* are  
961 reported in each part of the curves. Ticks on the curves are at f=0.9, 0.8, 0.7, 0.6, 0.5, 0.45, 0.4,  
962 0.35, 0.3, 0.25 and 0.2, starting from basanite RT-06I-354B (unreported for clarity). Symbols as  
963 in Fig. 2.

964 Supplementary Fig. S8: a) REE patterns of the Itasy mafic rocks ; b) the REE patterns of the  
965 mineral phases of RT-06I-355 and the host rock; c), d) REE patterns of the evolved rocks with c)  
966 model obtained by subtraction of 2% apatite and 1% titanite from sample RT-06I-387, by using  
967 the partition coefficients of Fedele et al. (2015) and Olin and Wolff (2012); and d) model  
968 obtained after subtraction of 0.7 and 1% titanite from sample RT-06I-387; e) The REE  
969 concentration of the Itasy rocks normalized to that of the most primitive sample (RT-06I-354B).  
970 The effects of titanite and apatite removal (cf. figure c) are also shown; f) Ratios between the

971 concentrations of the average composition of titanite, amphibole and clinopyroxene and that of  
972 the host rock. The experimental partition coefficients of titanite in phonolites (Olin and Wolff,  
973 2012), clinopyroxene in trachyphonolites (Fedele et al., 2009), and amphibole in trachytes  
974 (Tiepolo et al., 2007) are reported for comparison (data in Supplementary Table 11). g) The REE  
975 patterns of the Itasy titanites and those of alkaline intrusive rocks of within-plate affinity (data  
976 from Marks et al., 2008 and Melluso et al., 2010). Prometium is always interpolated. The  
977 chondrite is the CI of Boynton (1984).

Figure 1  
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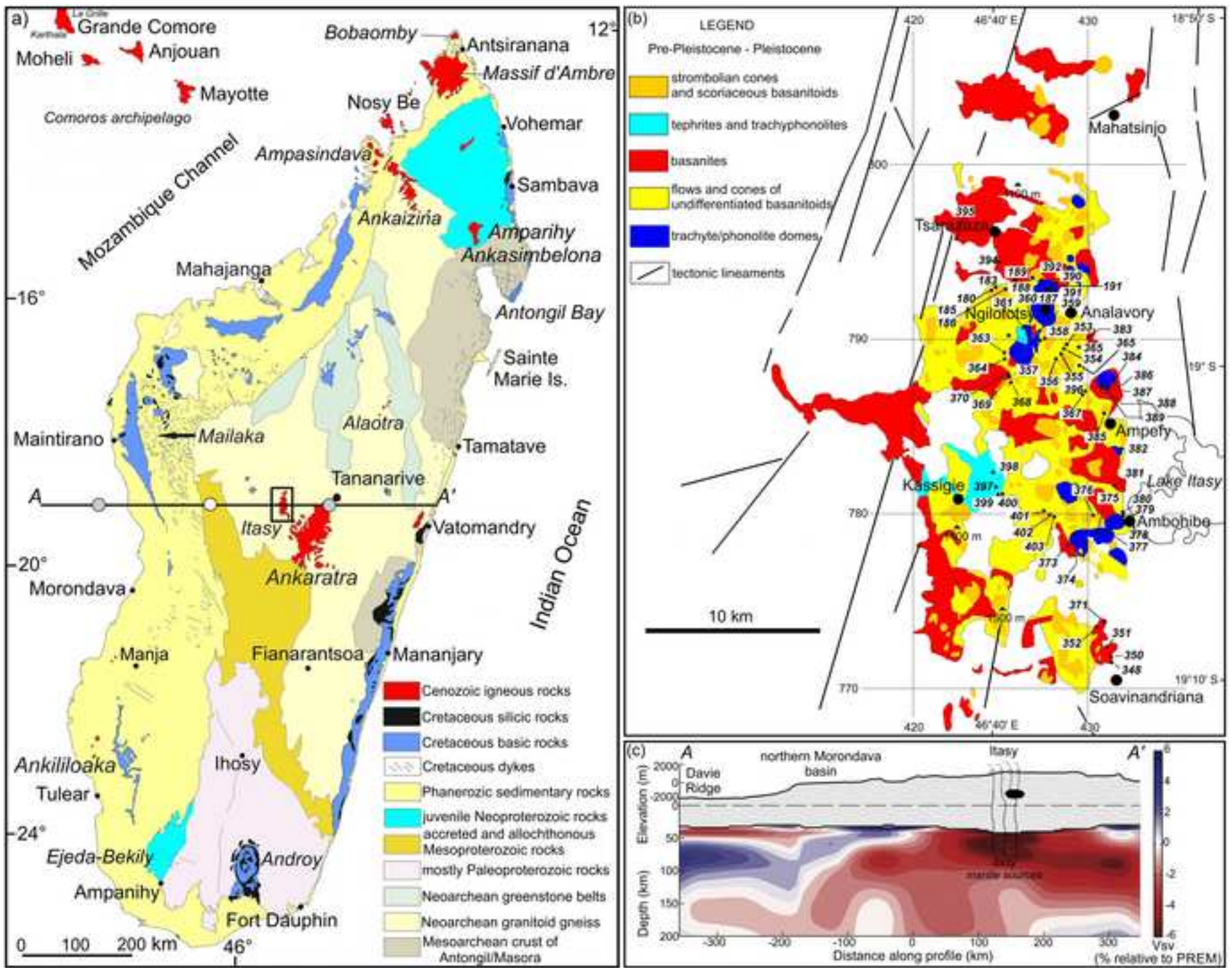




Figure 2

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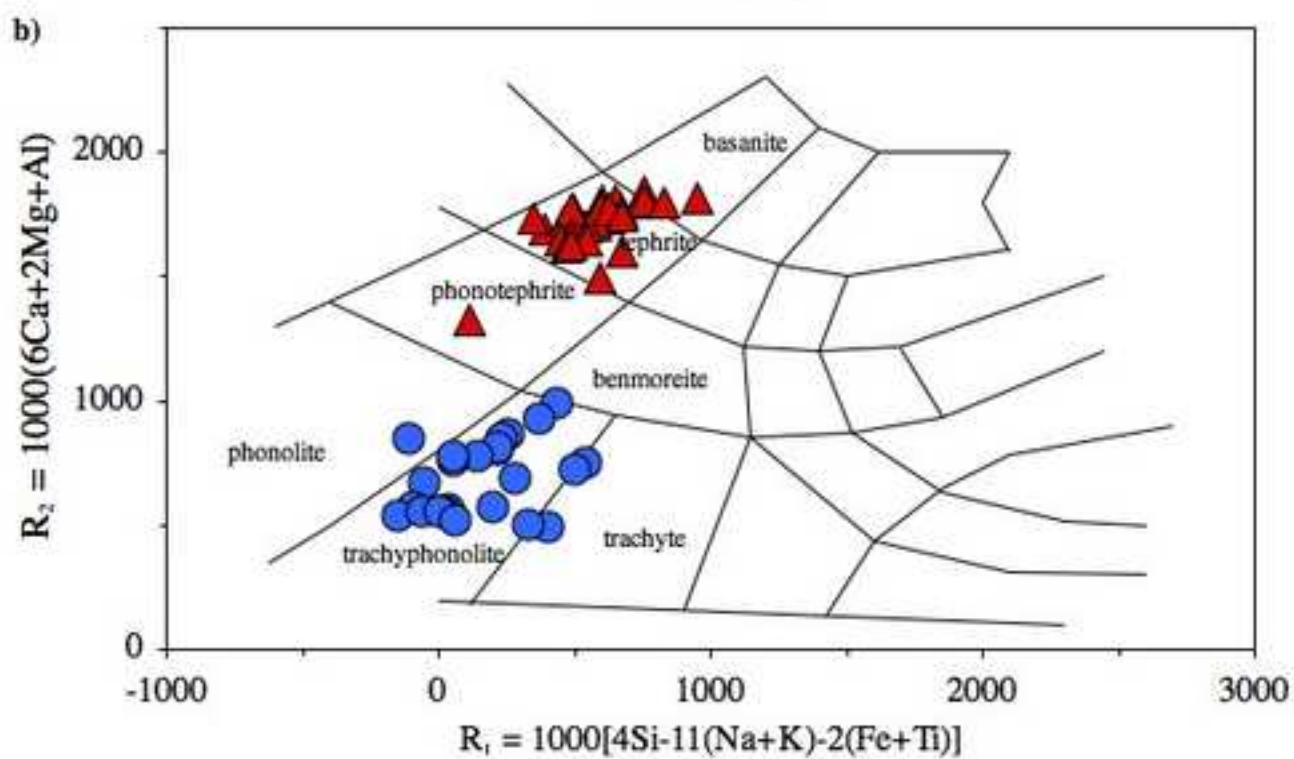
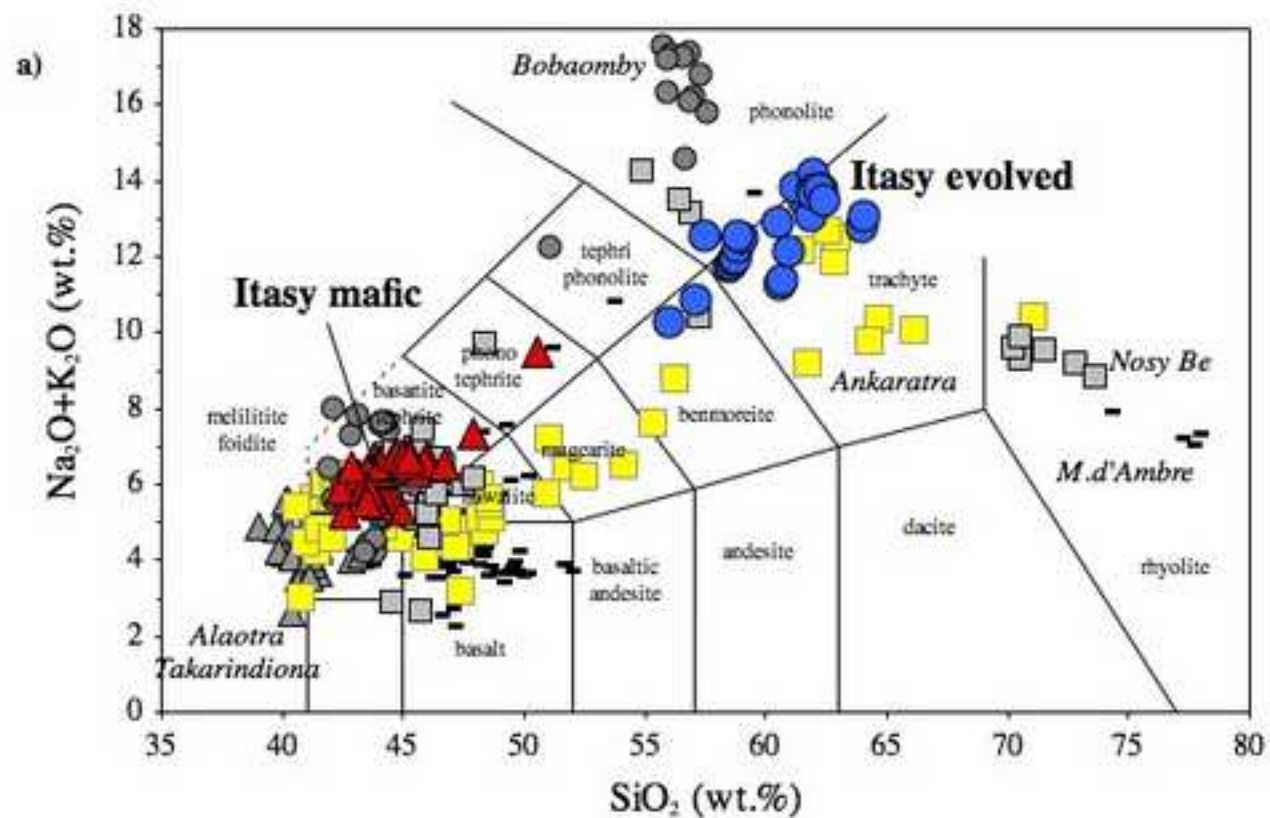


Figure 3

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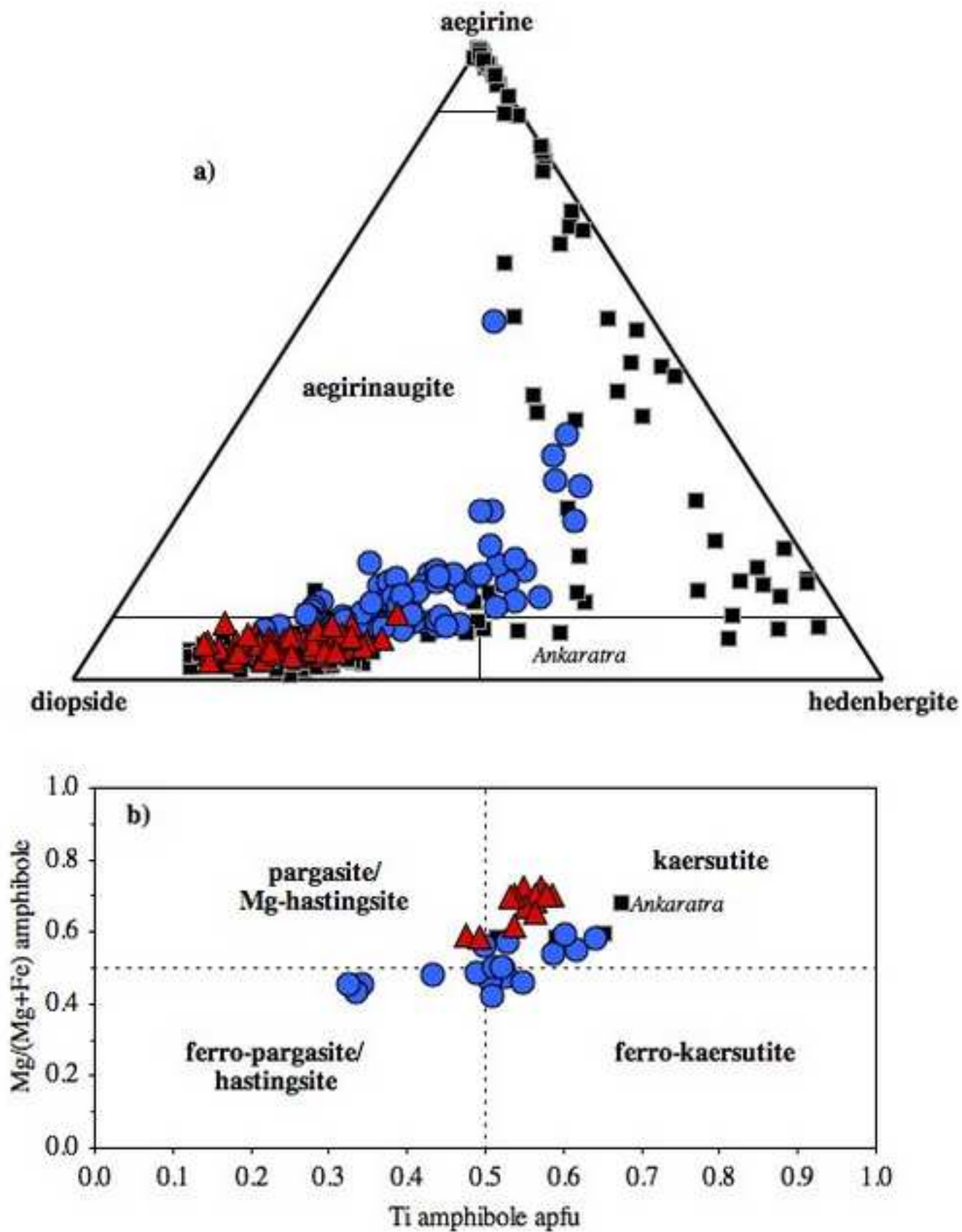


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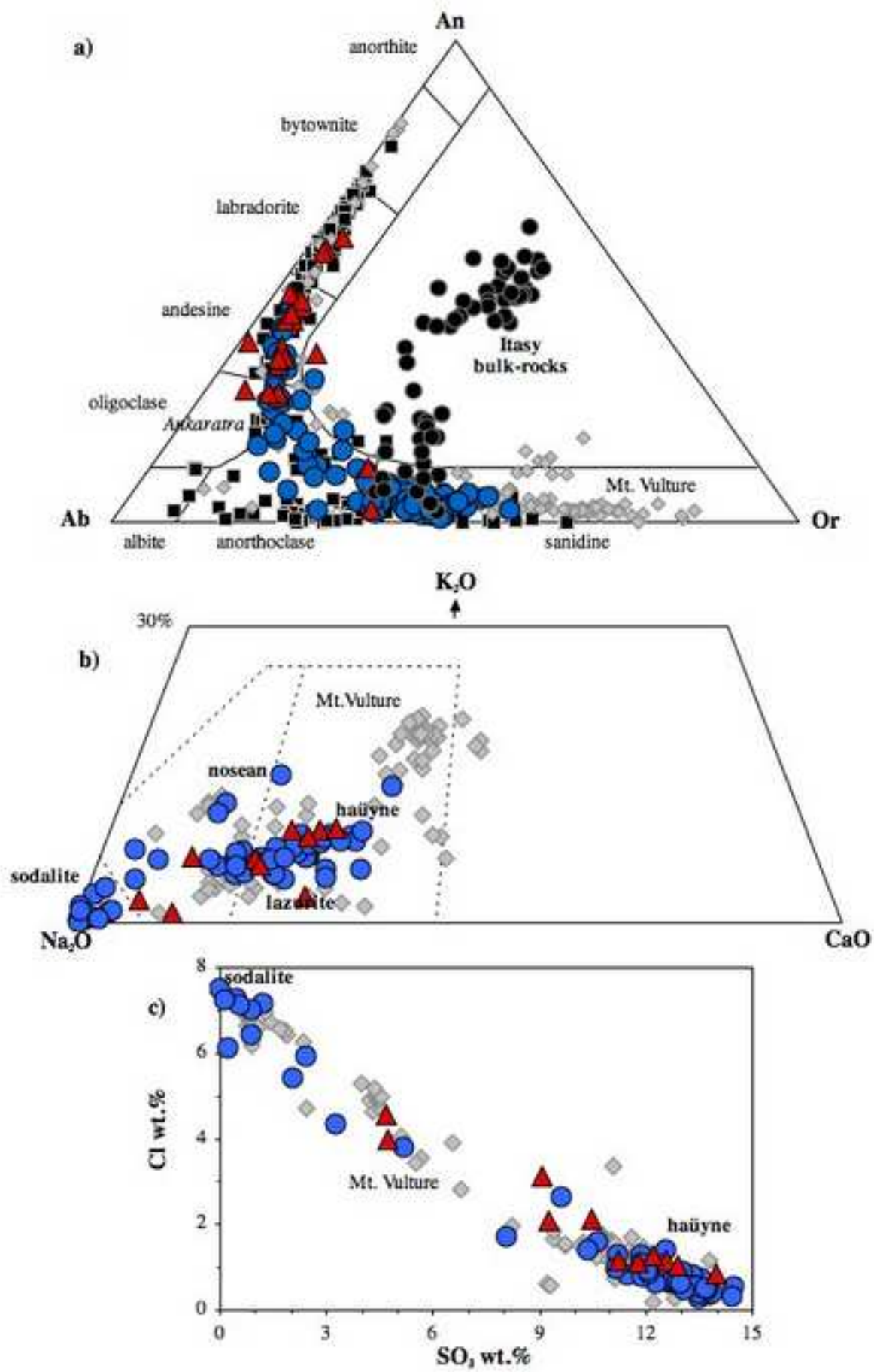


Figure 5

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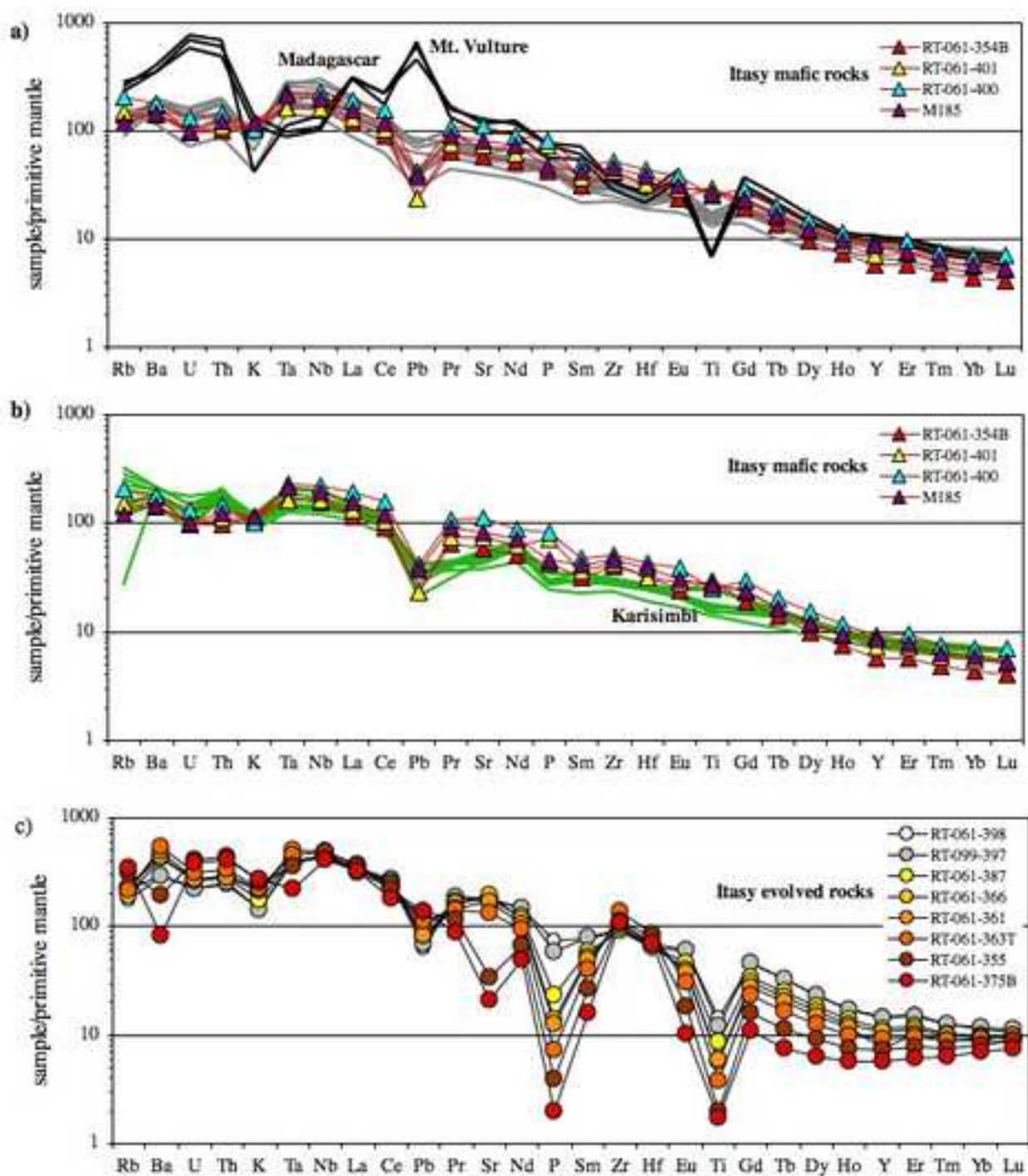


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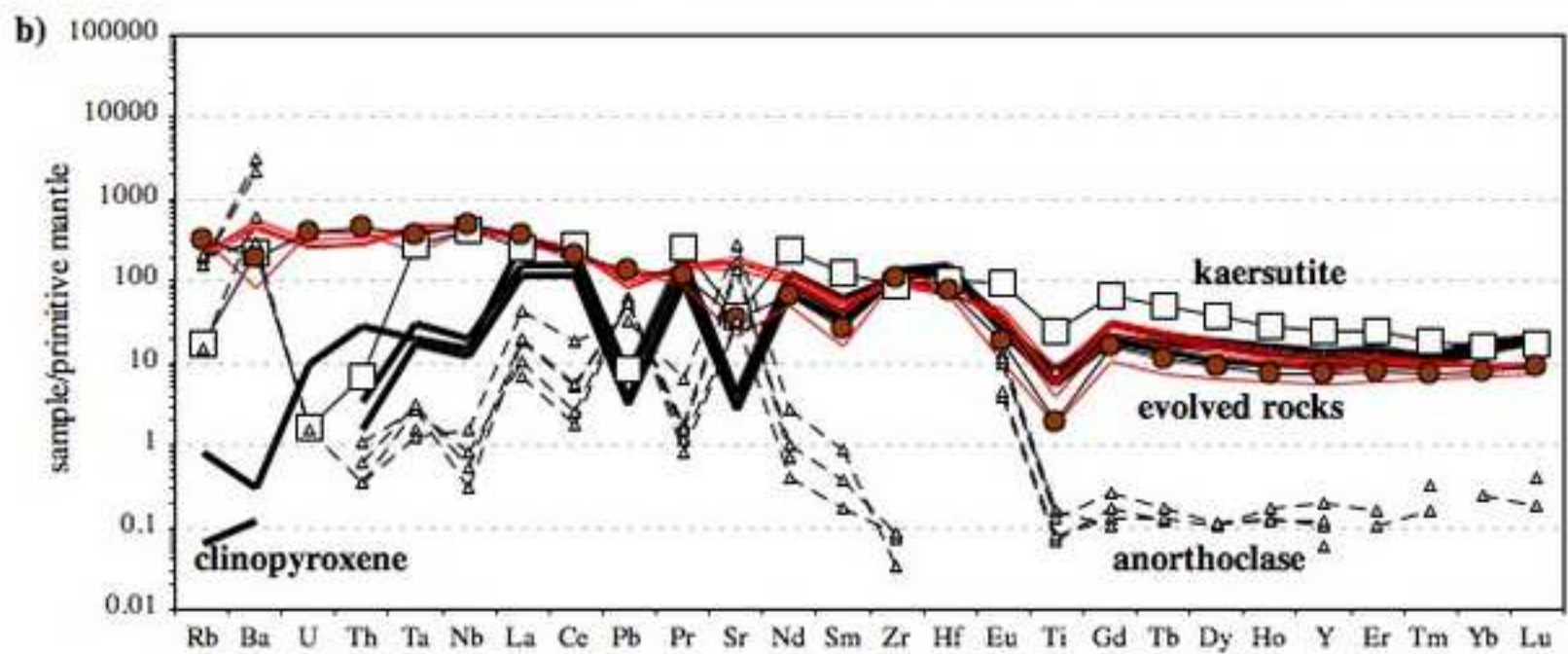
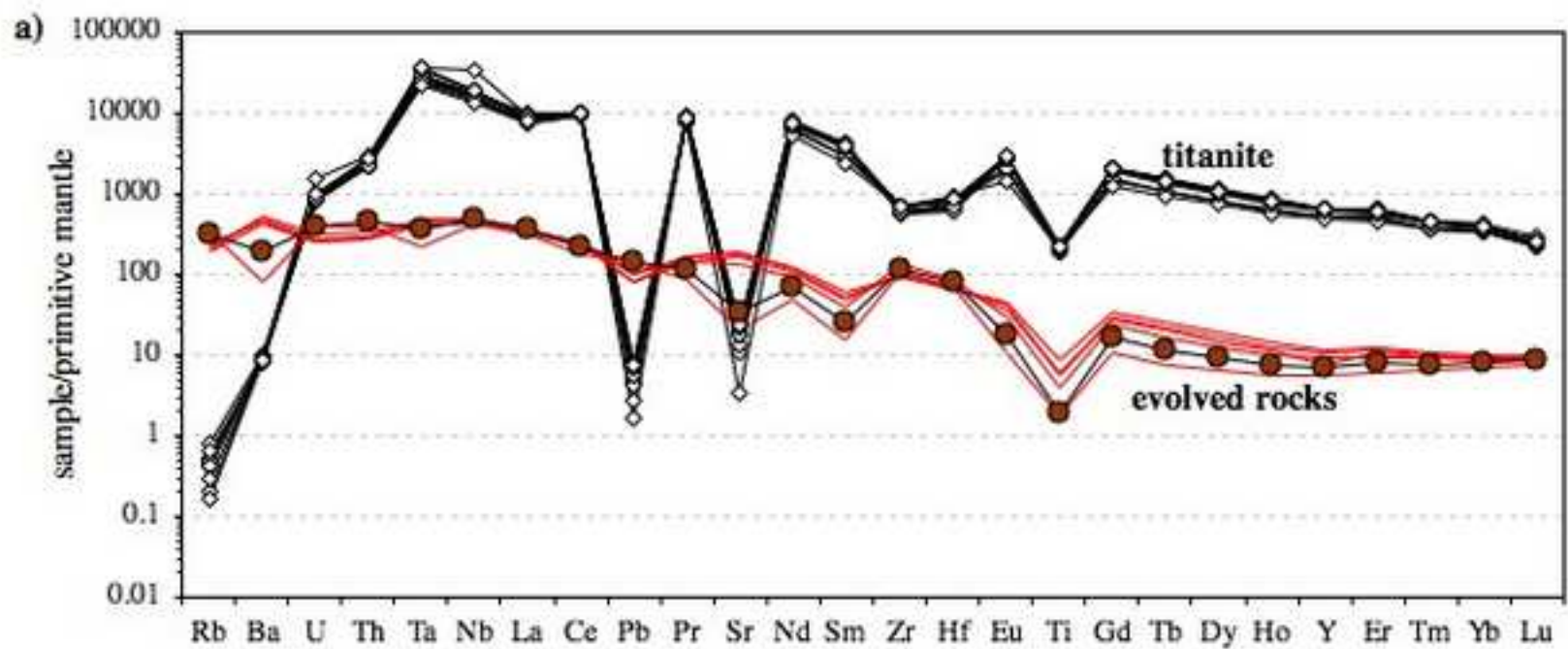


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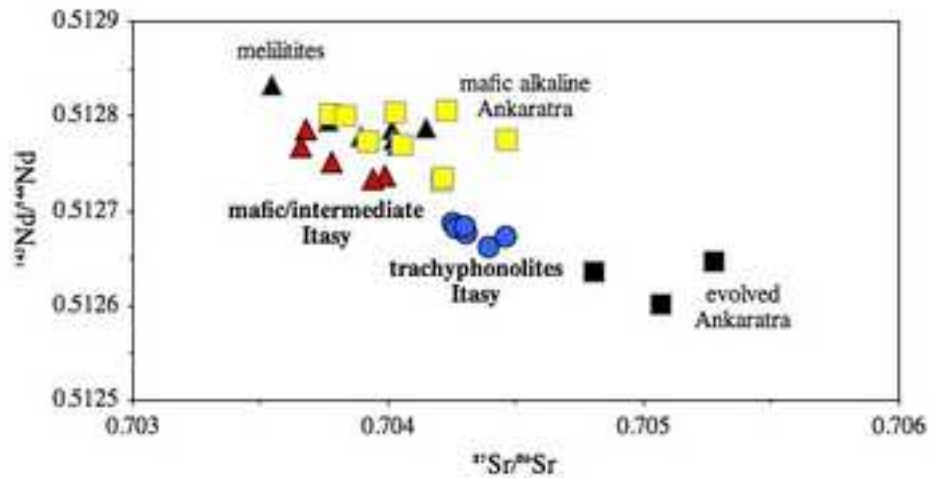
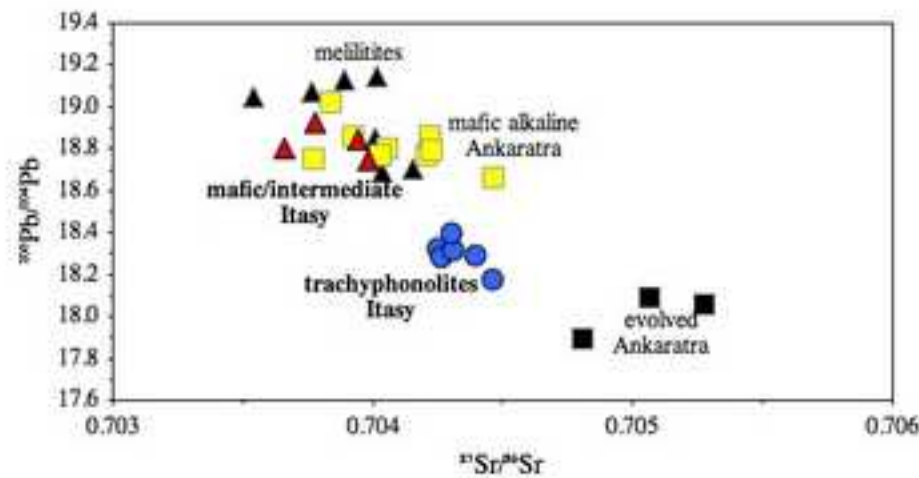
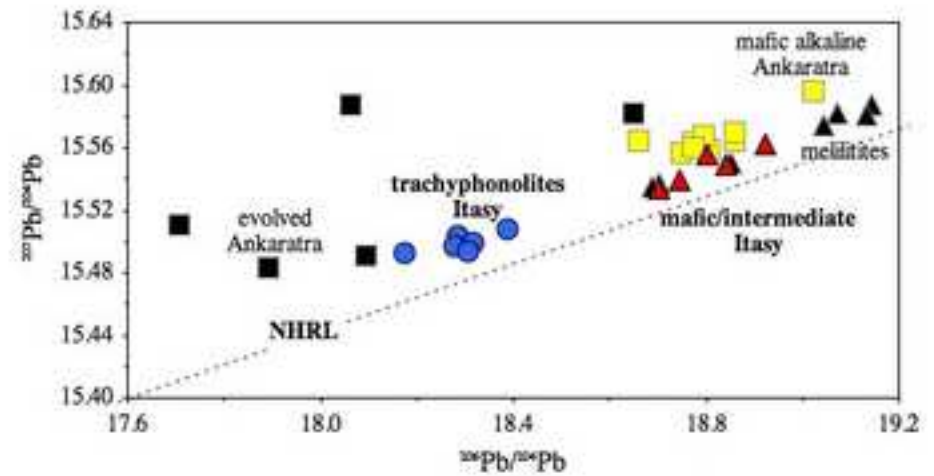
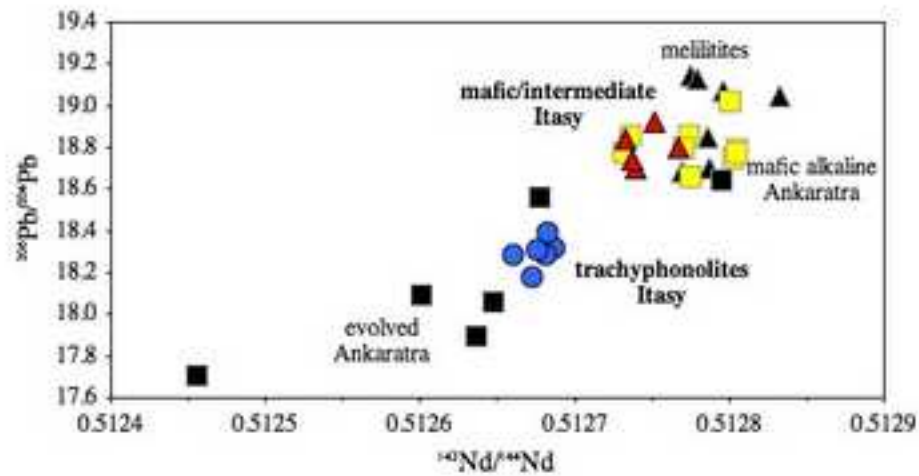


Figure 8  
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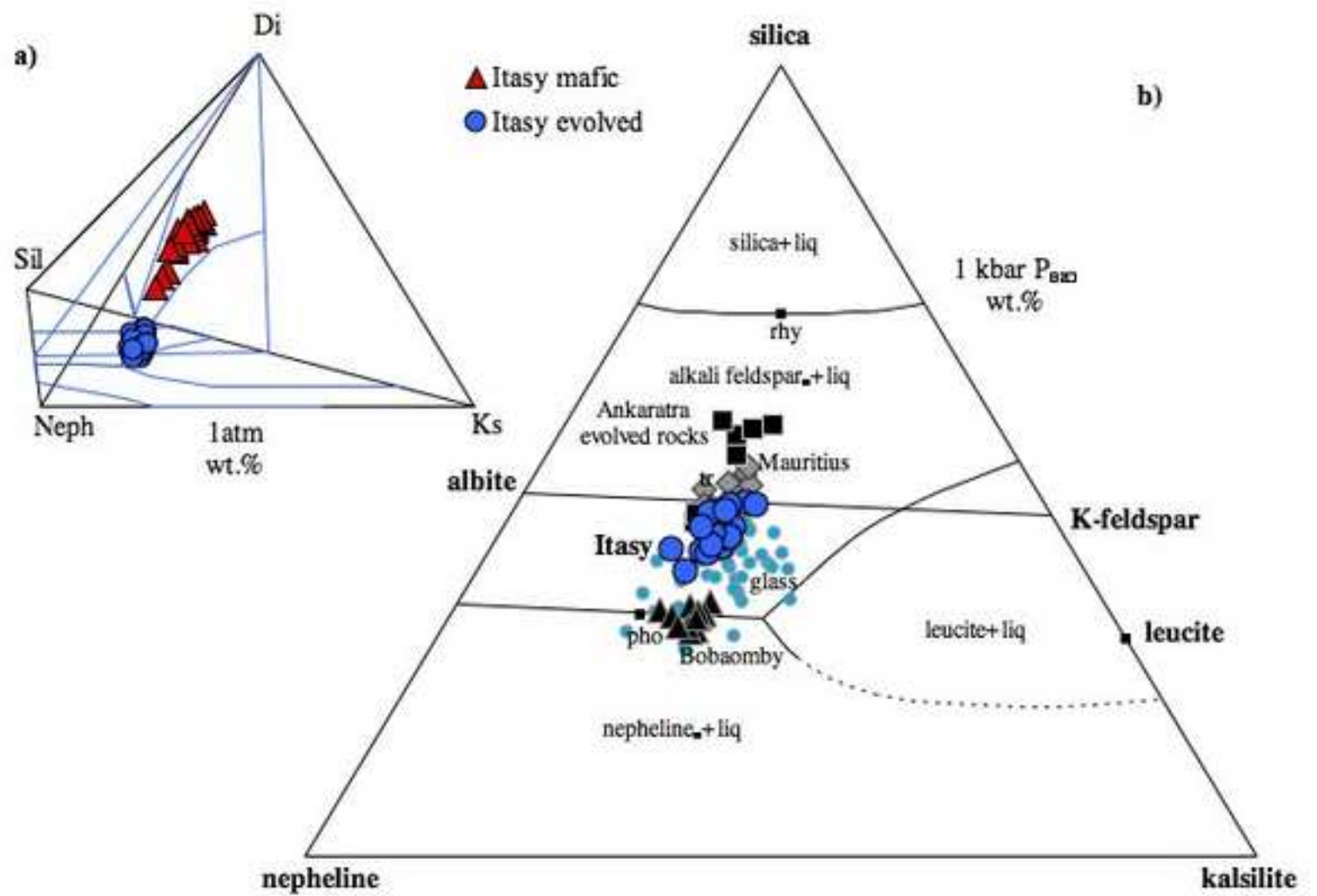


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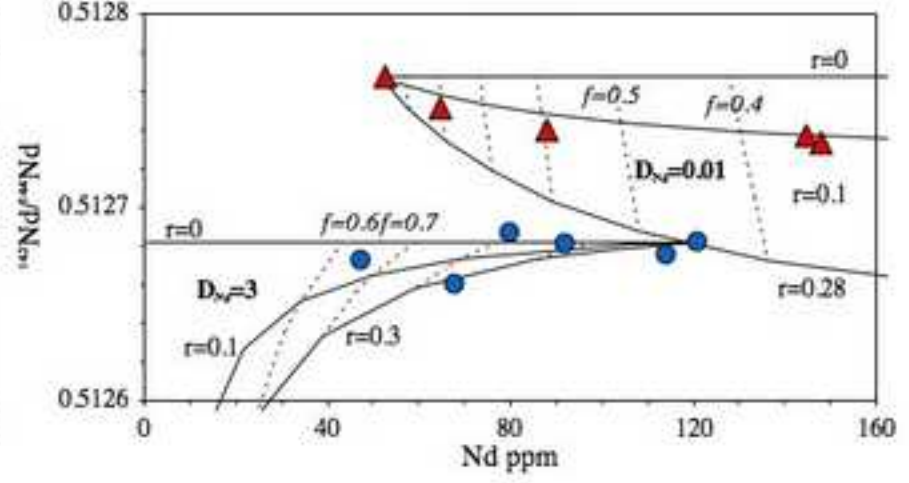
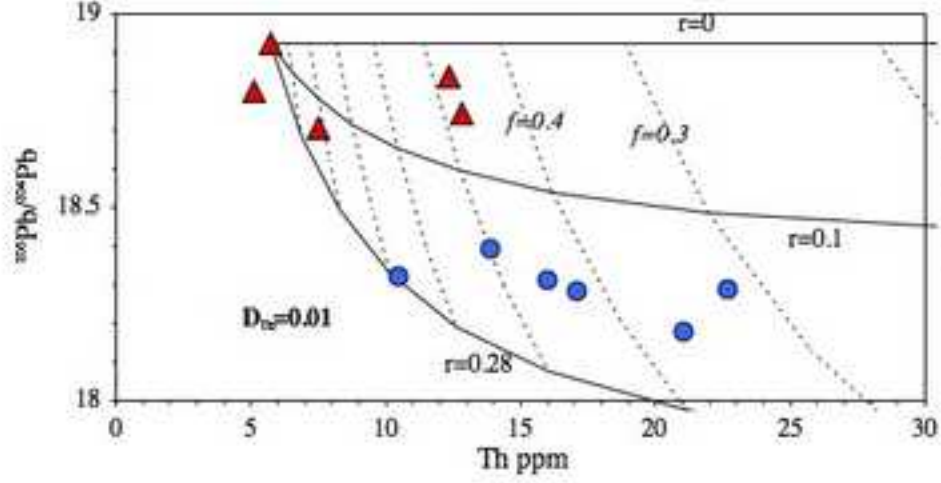
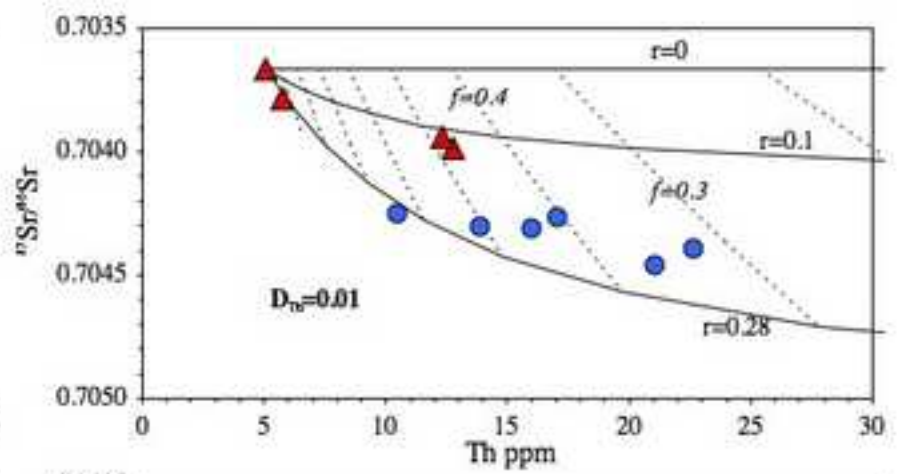
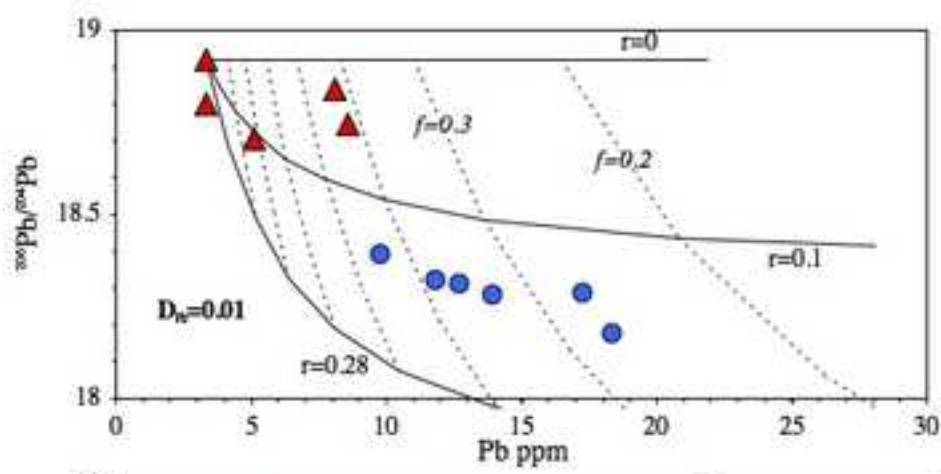




Table 1

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Table 1: ICP-MS major and trace element composition of Itasy rocks. A few element ratios are also reported

	BSN RT-061-354B	BSN M185	BSN RT-061-401	TPH RT-061-400	PHTPH RT-061-398	PHTPH RT-099-397	BEN RT-061-387	TRPHO RT-061-366	TRPHO RT-061-361	TRPHO RT-061-363T	TRPHO RT-061-355	TRPHO RT-061-375B
SiO <sub>2</sub>	41.75	46.23	40.94	43.06	49.17	48.94	54.26	57.45	57.97	59.92	60.90	61.00
TiO <sub>2</sub>	4.62	4.22	4.57	3.94	2.25	1.93	1.38	0.89	0.92	0.60	0.31	0.28
Al <sub>2</sub> O <sub>3</sub>	11.73	11.97	12.88	14.40	16.49	16.47	18.24	19.31	19.19	19.47	19.69	19.28
Fe <sub>2</sub> O <sub>3t</sub>	13.62	12.77	15.10	14.03	10.60	9.47	5.90	4.46	4.07	3.21	2.44	2.36
MnO	0.16	0.17	0.19	0.22	0.31	0.30	0.22	0.22	0.19	0.18	0.21	0.21
MgO	9.94	8.12	7.69	5.62	2.74	2.22	1.24	0.69	0.67	0.36	0.17	0.16
CaO	10.74	10.94	11.54	10.17	8.24	7.76	5.03	3.90	3.74	2.57	1.42	1.30
Na <sub>2</sub> O	2.42	2.18	2.63	3.56	5.54	4.69	5.49	6.80	6.35	5.60	6.56	7.42
K <sub>2</sub> O	2.64	2.69	2.44	2.34	3.37	3.23	3.99	4.83	4.83	5.08	5.90	6.17
P <sub>2</sub> O <sub>5</sub>	0.65	0.71	1.10	1.25	1.14	0.90	0.35	0.21	0.19	0.11	0.06	0.03
LOI	0.40	2.80	0.40	0.52	0.48	2.55	1.72	0.08	1.10	2.14	1.61	0.38
Total	98.66	102.80	99.48	99.11	100.3	98.46	97.83	98.84	99.23	99.25	99.27	98.59
Be	1	2	2	2	4	4	4	4	4	5	6	6
Sc	29		27	18	6	5	2	1	1			
V	407	353	395	301	126	104	76	50	47	29	17	16
Cr	330	184	200		80	50						80
Co	62	46	55	40	12	10	5	2	2	1		
Ni	170	98.7	80	30								
Cu	70	72.4	60	30								
Zn	110	155	120	160	220	220	160	160	140	150	180	160
Ga	23	25.6	24	27	32	31	31	31	30	30	35	32
Ge	2	1.18	2	2	2	2	2	2	2	1	2	2
Rb	65	55	69	94	81	102	89	107	104	96	142	157
Sr	918	1300	1170	1759	2776	2850	2727	3091	2666	2133	542	329
Y	19	29	24	31	49	48	38	36	25	31	24	19
Zr	341	387	362	434	781	804	768	816	870	1154	954	912
Nb	72	90	76	104	202	204	198	206	223	223	228	188
Mo		3.19	3	2	7	5	2	8	7	5	4	7
Ag	1.4	0.33	1.5	2.1	3.2	3.6	3.3	3.6	4.1		4.1	3.8
Sn	3	3.02	3	3	4	4	3	2	2	2	2	2
Cs		0.45	0.5	0.6	0.8	1.2	1	1	1	1.2	1.3	1.5
Ba	727	751	882	921	1482	1496	2135	2251	2468	2753	999	417
La	59.9	79.1	66.7	99.2	166.0	176.0	177.0	176.0	161.0	190.0	182.0	165.0
Ce	122.0	161.5	138.0	208.0	344.0	362.0	340.0	334.0	310.0	312.0	288.0	243.0
Pr	13.1	18.3	15.2	22.0	37.5	38.2	34.0	32.8	30.1	28.8	23.2	17.9
Nd	51.6	72.5	62.5	88.8	145.0	146.0	124.0	115.0	105.0	92.8	67.8	48.6
Sm	10.2	13.6	11.8	15.5	25.0	25.4	19.0	17.1	16.1	12.9	8.5	5.1
Eu	2.9	3.8	3.6	4.7	7.4	7.5	5.7	5.0	4.6	3.7	2.2	1.3
Gd	8.3	10.2	10.0	12.7	20.2	19.6	14.6	13.0	12.1	10.2	6.9	4.7
Tb	1.1	1.3	1.3	1.6	2.6	2.6	2.0	1.8	1.6	1.3	0.9	0.6
Dy	5.3	6.7	6.4	8.2	12.7	12.5	10.0	8.8	8.0	6.8	5.0	3.4
Ho	0.90	1.16	1.10	1.40	2.10	2.10	1.70	1.50	1.40	1.20	0.90	0.70
Er	2.00	2.65	2.50	3.30	5.30	5.10	4.20	3.90	3.60	3.30	2.70	2.10
Tm	0.26	0.34	0.32	0.41	0.67	0.67	0.56	0.54	0.49	0.49	0.41	0.34
Yb	1.50	2.03	1.90	2.40	3.90	4.00	3.40	3.30	3.10	3.20	2.80	2.40
Lu	0.22	0.29	0.28	0.38	0.57	0.61	0.50	0.52	0.47	0.54	0.48	0.40
Hf	7.4	8.7	7.2	9.8	14.3	15.6	14.7	14.7	16	19.2	18	15.5
Ta	5.1	6.51	4.9	7	11.8	12	12.6	12	15.1	13.6	11	6.8
Pb	3.35	5.44	3.35	6	9	10	12	14	12	16	20	20
Th	6.2	7.7	6.9	9.2	14.7	15.9	17.7	18.9	17.8	21.2	27.6	25.3
U	1.7	1.75	1.7	2.3	4	3.8	4.5	4.5	4.7	5.4	7	6.7
Zr/Hf	46	44	50	44	55	52	52	55	54	60	53	59
Nb/Ta	14	14	15	15	17	17	16	17	15	16	21	28
Th/U	3.6	4.4	4.1	4.0	3.7	4.2	3.9	4.2	3.8	3.9	3.9	3.8
La/Yb <sub>n</sub>	26.9	26.3	23.7	27.9	28.7	29.7	35.1	36.0	35.0	40.0	43.8	46.4

BSN, basanite; TPH, tephrite; PHTPH, phonotephrite; BEN, benmoreite; TRPHO, trachyphonolite; La/Yb<sub>n</sub> is the chondrite normalized ratio

Table 2

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Table 2: LAM-ICP-MS and microprobe composition of titanite, amphibole, clinopyroxene and feldspar in the trachyphonolite RT061355

Table 2: LAN

ppm	RT-061-355	RT-061-355	RT-061-355	RT-061-355	RT-061-355	RT-061-355	RT-061-355	RT-061-355	RT-061-355	RT-061-355	RT-061-355	RT-061-355	RT-061-355	RT-061-355
	ttn	ttn	ttn	ttn	ttn	ttn	ttn	ttn	ttn	ttn	ttn	ttn	ttn	av ttn
Li	0.54	0.76	2.02	2.23	2.01	0.63								1.4
B	47.78	49.26	43.04	69.42	73.42	68.24		72.75	74.78	69.97	52.65			62.1
Sc	7.70	7.92	8.31	8.05	8.12	7.58		5.27	5.36	6.42	5.79			7.1
Ti	178310	178699	181535	204264	206175	192850		212573	207348	207720	211055			198053
V	305.8	324.9	324.1	286.0	311.4	235.7		326.4	318.9	327.6	316.2			307.7
Cr	2.65													2.7
Co	0.13	0.12	0.11	0.34				0.31	0.19					0.2
Ni	0.52	0.44	0.71		0.42			0.95	1.39	0.87	0.85			0.8
Zn	50.97	47.43	50.99	37.29	47.25	48.41		27.12	25.45	26.98	30.48			39.2
Rb	0.20	0.20	0.24	0.36	0.23	0.19		0.10	0.08	0.30	0.13			0.2
Sr	174.13	291.71	343.76	204.66	276.91	54.22		348.77	275.16	359.25	382.21			271.1
Y	1790.1	2249.9	2146.0	1734.8	2188.4	1616.5		2199.4	2223.5	2054.2	2198.4			2040.1
Zr	5834.7	4930.9	5907.6	5933.6	5035.2	5702.6		4830.0	4574.9	5043.0	5836.6			5362.9
Nb	8629.9	7681.5	8876.7	8424.2	7207.8	15230.2		6469.2	6949.3	6168.9	8601.3			8423.9
Cs	0.02	0.01	0.01			0.04								0.02
Ba	46.10	47.03	48.57	49.06	46.57	48.21		43.47	45.29	42.16	45.08			46.2
La	4557.5	4600.0	4542.1	5055.2	4224.2	4890.9		3922.8	4056.2	3840.3	4142.6			4383.2
Ce	13139.3	13550.0	13500.0	13786.1	13157.6	13121.1		12700.0	13178.3	12082.8	12951.9			13116.7
Pr	1593.1	1816.5	1837.8	1653.2	1777.0	1483.3		1749.4	1769.1	1629.9	1754.4			1706.4
Nd	6030.2	7720.2	7733.4	6332.8	7680.5	5332.0		7605.9	7727.5	7047.1	7614.2			7082.4
Sm	922.3	1318.5	1306.6	949.3	1315.2	781.8		1351.1	1338.1	1218.3	1293.2			1179.4
Eu	236.1	344.3	339.9	252.5	345.5	179.6		362.6	349.0	332.1	351.7			309.3
Gd	618.4	888.5	873.2	644.6	883.3	520.9		897.9	874.7	806.5	868.6			787.7
Tb	85.0	121.5	116.9	88.2	118.6	75.1		123.4	120.8	110.0	118.5			107.8
Dy	424.6	596.9	583.6	441.0	598.4	392.0		614.1	600.1	553.1	587.2			539.1
Ho	73.4	99.4	97.4	77.3	100.5	68.7		101.8	104.0	92.0	98.2			91.3
Er	171.4	218.5	215.2	174.8	226.2	160.1		216.4	217.8	199.9	213.9			201.4
Tm	19.9	24.8	24.4	19.9	25.3	19.5		24.5	25.0	23.1	25.0			23.1
Yb	116.1	135.9	134.4	119.8	141.3	117.6		136.6	141.1	127.6	133.0			130.3
Lu	12.1	13.8	13.7	12.7	15.4	11.9		13.7	13.2	13.6	14.0			13.4
Hf	183.1	155.1	189.4	201.9	171.4	211.4		151.2	140.5	156.6	191.8			175.2
Ta	718.8	838.6	1047.6	866.5	865.9	1133.7		708.5	763.4	655.2	1102.7			870.1
Pb	0.87	0.65	0.73	0.62	0.40	1.16		0.24	0.79	0.98	1.07			0.8
Th	151.6	137.8	154.8	168.0	140.6	187.3		133.6	141.4	139.6	173.8			152.9
U	17.80	14.73	15.67	16.48	14.39	26.76		15.03	15.11	15.45	16.74			16.8
SREE	27999	31449	31319	29607	30609	27154		29820	30515	28076	30166			SREE
Zr/Hf	31.9	31.8	31.2	29.4	29.4	27.0		31.9	32.6	32.2	30.4			Zr/Hf
Nb/Ta	12.0	9.2	8.5	9.7	8.3	13.4		9.1	9.1	9.4	7.8			Nb/Ta
Zr/Y	3.3	2.2	2.8	3.4	2.3	3.5		2.2	2.1	2.5	2.7			Zr/Y
Th/U	8.5	9.4	9.9	10.2	9.8	7.0		8.9	9.4	9.0	10.4			Th/U
La/Ybn	26.5	22.8	22.8	28.4	20.2	28.0		19.4	19.4	20.3	21.0			La/Ybn

**Table 3**[Click here to download Table: Table3itasy.xls](#)

Table 3: Isotopic composition of the Itasy rocks. The composition of the standards which were run with the unknowns is also reported

		$^{143}\text{Nd}/^{144}\text{Nd}$	$\pm 2\text{s internal} * 10^6$	$^{87}\text{Sr}/^{86}\text{Sr}$	$\pm 2\text{s internal} * 10^6$	$^{206}\text{Pb}/^{204}\text{Pb}$	$\pm 2\text{s internal}$	$^{207}\text{Pb}/^{204}\text{Pb}$	$\pm 2\text{s internal}$	$^{20}\text{Pb}/^{204}\text{Pb}$	$\pm 2\text{s internal}$		
BSN	RT-06I-354B	0.512768	10	0.703661	10	18.801	0.0010	15.556	0.0010	39.012	0.0027		
BSN	RT-06I-400	0.512740	15			18.704	0.0009	15.534	0.0008	38.891	0.0027		
BSN	IT185	0.512786	10	0.703678	10								
BSN	RT-06I-401	0.512752	12	0.703780	11	18.922	0.0008	15.562	0.0008	39.088	0.0027		
phtph	RT-06I-397	0.512737	14	0.703984	12	18.744	0.0008	15.539	0.0007	38.965	0.0022		
phtph	RT-06I-398	0.512733	11	0.703942	10	18.840	0.0008	15.549	0.0008	39.027	0.0021		
ben	RT-06I-387	0.512682	9	0.704303	10	18.390	0.0009	15.507	0.0009	38.567	0.0025		
trph	RT-06I-355	0.512660	7	0.704397	12	18.285	0.0009	15.503	0.0009	38.486	0.0029		
trph	RT-06I-361	0.512687	13	0.704253	10	18.318	0.0010	15.499	0.0008	38.494	0.0026		
trph	RT-06I-363T	0.512681	11	0.704266	11	18.279	0.0007	15.496	0.0008	38.464	0.0024		
trph	RT-06I-366	0.512676	7	0.704312	11	18.308	0.0010	15.493	0.0010	38.493	0.0029		
trph	RT-06I-375B	0.512673	13	0.704462	13	18.176	0.0008	15.492	0.0009	38.380	0.0000		
Std	BHVO-2	0.512985	8 (0.512984 $\pm$ 11, Weis et al., 2006)										
Std	BHVO-2	0.512974	11 (0.512987 $\pm$ 19, long-term UCT average n=120/exclude 1)										
Std	BHVO-2	0.512976	9 (0.512980 $\pm$ 12, GEOREM)										
Std	ref JNdi-1	0.512115	7 (Tanaka et al., 2000)										
Std	BHVO-2			0.703488	12 (0.703479 $\pm$ 20, Weis et al., 2006)								
Std	BHVO-2			0.703452	13 (0.703489 $\pm$ 44, long-term UCT average n=124/exclude 6)								
Std	BHVO-2			0.703466	10 (0.703469 $\pm$ 17, GEOREM)								
Std	BHVO-2			0.703467	15								
Std	BHVO-2					18.6506	0.0007	15.5358	0.0006	38.2332	0.0019		
Std	BHVO-2					18.6920	0.0008	15.5376	0.0010	38.2597	0.0028		
Std	BHVO-2						Weis et al., 2006	18.6474	0.0242	15.5334	0.0094	38.2367	0.0182
Std	BHVO-2						GEOREM range	18.514-18.687		15.457-15.558		37.992-38.294	
Std	BHVO-2						(long-term UCT average n=58/exclude 3)	18.6306	0.0620	15.5337	0.0139	38.2237	0.0483
Std	ref NIST981						Galer & Abouchami 1998	16.9405	0.0015	15.4963	0.0016	36.7219	0.0044

BSN, basanite; phtph, phonotephrite; ben, benmoreite; trph, trachyphonolite

supplfig1

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supplfig1b

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supplfig1c

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supplfig2

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supplfig3

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